

Chapter 9

Celestial Observations and Sight Reduction Methods

Introduction

In this chapter you will learn the methods that are necessary to complete a day's work in navigation. These include several ways of finding gyrocompass error, reducing sunlines and moonlines, finding latitude by LAN or Polaris, and reducing sights of stars and planets.

We've already covered the basics of celestial navigation in chapter 6. You may want to occasionally refer back to that material to have a clearer understanding of this material. We will discuss the procedure aspect of performing and reducing celestial observations only. As you become more familiar with this subject, you are advised to increase your knowledge by studying references such as *Dutton's Navigation and Piloting* and *Bowditch* Pub No. 9.

Objectives

The material in this chapter will enable the student to:

- Determine gyrocompass error by azimuth of the Sun and Polaris, and amplitude of the Sun.
- Reduce sights taken on the Sun, Moon, stars, and planets using H.O. 229, *Sight Reduction Tables for Marine Navigation*, and the *Nautical Almanac*.
- Reduce sights taken on the stars using H.O. 249, *Sight Reduction Tables for Air Navigation* and the *Air Almanac*.
- Determining latitude by local apparent noon.
- Plot celestial LOPs based on assumed positions.

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Methods For Finding Gyrocompass Error

Introduction

There are three celestial methods used by QMs for finding gyrocompass error. They are:

- Azimuth of the Sun
- Azimuth of Polaris
- Amplitude of the Sun

In each case, you are required to gather data for use in computation. This data may be in the form of sights from the sextant, time in GMT, DR Lat. and Long., and so forth. For each celestial method, we will begin with gathering the necessary data and then working the solutions.

Azimuth of the Sun

You must know the following values to determine gyrocompass error by azimuth:

Gathering Data

- Time of the actual observation
- Date of the observation
- DR position at the time of observation
- Azimuth (gyro bearing of the Sun)

Rule: Due to the elevation of the Sun, azimuths should be taken in mid-morning or mid-afternoon.

Use the following table to gather the data to work the azimuth solution. You must have a recorder present to mark and record the exact time of the observation

Step	Action
1.	Obtain a time tick from WRN-6 or chronometer with a stopwatch.
2.	Break out and place the azimuth circle on the gyro repeater closest to the Sun.
3.	Align the Sun in the reflecting mirror in a manner so that the rays reflect back through the prism housing and onto the compass card.
4.	When each spirit level is leveled, mark the time and record the reflected gyro bearing from the compass card to the nearest 0.1°. Note: This is a difficult procedure in heavy seas; however, if the azimuth circle is not level, errors will occur.
5.	Repeat steps 4 and 5 a minimum of three times.

Azimuth of the Sun, Continued

Gathering Data Now that we have three good observations, we need only to find the DR position for each observation to have the data we need to find the azimuth of the Sun. We will work an example problem using OPNAV strip form AZIMUTH BY PUB 229. For brevity, we'll work on one observation only. In actual practice, it's faster to work out all three at the same time by placing the strip form on the left and working the three observations in the next three columns. The purpose of taking at least three observations is to allow us to find errors when taking observations and averaging gyro error. This process normally provides the best results in determining total gyro error.

Besides the data from the observation, you'll need the *Nautical Almanac* and Pub 229 to solve the problem.

From the strip form and publication we will find out exactly what the gyro bearing *should* read, then we will compare that value to the actual gyro bearing from the observation. The result will be our gyro error.

Example Problem

For our example problem, we will assume that we have gathered the following data:

Date: 19 Nov 84
DR Lat.: $33^{\circ} 37' \text{ N}$
DR Long.: $112^{\circ} 39' \text{ E}$
ZT: 15h 42m 22s
Gyro Brg: 231.6°

On the following pages, you will find the page laid out with the blank strip form on the left, the action steps in the middle, and the result on the right.

Azimuth of the Sun, Continued

OPNAV 3130/ Azimuth by 229	ACTION	Completed Strip Form
Date:	Enter the date.	19 NOV 84
DR Posit	Enter the DR position.	33°37'N - 112°39'E
Body	Enter the name of the body observed.	Sun
GMT	Enter the time GMT.	07h42m22s
GHA(h)	Enter the GHA hour value from the <i>Nautical Almanac</i> (fig. 9-1)	288° 38.9'
Increment (m/s)	Enter the minutes and seconds value from the <i>Nautical Almanac</i> (fig. 9-2).	10° 35.5'
Total GHA	Add GHA(h) and increments (m/s).	299° 14.4'
DR Long +E, -W (+ - 360° if needed)	Enter the DR Longitude, add east or subtract west.	112° 39.0'E
LHA	LHA= Total GHA +E or -W DR Long.	51° 53.4'
Tab Dec	Enter the tabulated declination for 07 hours on the Sun column from the <i>Nautical Almanac</i> .	S 19° 31.2'
d# / D Corr ⁿ	<p>The d# is found at the bottom of the Sun Dec column, in this case it is +0.6. It is assigned a + because Dec is increasing (0700= 19°31.2 0800= 19°31.8). You MUST assign a + or - to the d#.</p> <p>The D corr is found on the Increments and Corrections page for 42m 22s. Look under the <i>v</i> or <i>d</i> column for the d# (0.6) and record the Corrⁿ value (0.4). The D Corrⁿ assumes the same sign as the d# .</p>	+0.6 / +0.4
True Dec	Apply the D Corr ⁿ to Tab Dec	S 19° 31.6
DR Lat same or contrary	Enter the whole degree of latitude and determine if it is named (N or S) as True Dec In this case, Lat. is N and Dec is S. so it is contrary.	N 33° contrary

Up to this point, we have worked the strip form to obtain three values, LHA, True Dec., and DR Latitude. We now have everything we need to enter Pub 229. Pub 229 is entered using whole degrees of Lat., LHA, and Dec. only. We will also interpolate the leftover values using Pub 229.

Azimuth of the Sun, Continued

1984 NOVEMBER 17, 18, 19 (SAT., SUN., MON.)															225	
G.M.T.	SUN				MOON				Lat.	Twilight		Sunrise	Moonrise			
	G.H.A.	Dec.	G.H.A.	U	Dec.	d	H.P.	Naut.		Civil	17		18	19	20	
17 00	183 46.2	S18 58.5	258 20.6	9.9	N14 55.5	13.4	59.0	N 72	07 11	08 49	10 12	23 36	26 02	02 02	04 25	
01	198 46.1	59.1	272 49.5	9.9	14 42.1	13.5	59.1	N 70	06 59	08 24	10 12	23 51	26 05	02 05	04 17	
02	213 46.0	18 59.7	287 18.4	10.0	14 28.6	13.6	59.1	68	06 50	08 04	09 28	24 03	26 03	02 07	04 11	
03	228 45.8	19 00.3	301 47.4	10.0	14 15.0	13.6	59.1	66	06 42	07 49	08 59	24 13	26 03	02 09	04 06	
04	243 45.7	00.9	316 16.4	10.1	14 01.4	13.7	59.1	64	06 35	07 36	08 37	24 21	26 03	02 11	04 02	
05	258 45.6	01.6	330 45.5	10.0	13 47.7	13.9	59.2	62	06 29	07 25	08 20	24 28	26 02	02 13	03 58	
06	273 45.5	S19 02.2	345 14.5	10.2	N13 33.8	13.9	59.2	60	06 23	07 16	08 06	24 34	26 02	02 14	03 55	
07	288 45.4	02.8	359 43.7	10.1	13 19.9	14.0	59.2	N 58	06 18	07 07	07 54	24 39	26 01	02 15	03 52	
S 08	303 45.2	03.4	14 12.8	10.2	13 05.9	14.1	59.2	56	06 14	07 00	07 43	24 44	26 00	02 16	03 50	
A 09	318 45.1	04.0	20 42.0	10.2	12 51.8	14.1	59.3	54	06 10	06 53	07 34	24 48	26 00	02 17	03 48	
T 10	333 45.0	04.6	26 42.0	10.2	12 37.7	14.3	59.3	52	06 06	06 47	07 25	24 52	26 00	02 18	03 46	
U 11	348 44.8	05.2	32 40.5	10.3	12 23.4	14.3	59.3	50	06 02	06 42	07 18	24 55	26 00	02 19	03 44	
R 12	3 44.7	S19 05.8	38 40.5	10.3	12 09.1	14.4	59.3	45	05 54	06 30	07 02	25 03	26 00	02 21	03 40	
O 13	18 44.6	06.4	44 36.7	10.3	11 54.7	14.5	59.3	N 40	05 47	06 19	06 49	25 09	26 00	02 22	03 37	
A 14	33 44.5	07.0	50 36.7	10.4	11 40.2	14.5	59.4	35	05 40	06 10	06 37	25 15	26 00	02 23	03 34	
Y 15	48 44.3	07.6	56 36.7	10.4	11 25.7	14.6	59.4	30	05 33	06 02	06 28	25 21	26 00	02 25	03 31	
16	63 44.2	08.2	62 36.7	10.5	11 11.1	14.7	59.4	20	05 21	05 47	06 11	25 26	26 00	02 26	03 27	
17	78 44.1	08.8	68 36.7	10.5	10 56.4	14.8	59.4	N 10	05 08	05 34	05 56	25 31	26 00	02 28	03 23	
18	93 44.0	S19 09.4	74 36.7	10.5	10 41.6	14.8	59.5	0	04 54	05 20	05 42	25 36	26 00	02 30	03 20	
19	108 43.8	10.0	80 36.7	10.5	10 26.8	14.9	59.5	S 10	04 39	05 05	05 27	25 41	26 00	02 31	03 17	
20	123 43.7	10.6	86 36.7	10.5	10 11.9	15.0	59.5	20	04 20	04 48	05 12	25 46	26 00	02 33	03 13	
21	138 43.6	11.2	92 36.7	10.6	9 56.9	15.0	59.5	30	03 57	04 28	04 54	25 51	26 00	02 35	03 09	
22	153 43.4	11.8	98 36.7	10.6	9 41.9	15.1	59.5	35	03 42	04 16	04 44	25 56	26 00	02 36	03 07	
23	168 43.3	12.4	104 36.7	10.6	9 26.8	15.2	59.6	40	03 23	04 01	04 32	26 01	26 00	02 38	03 05	
18 00	183 43.2	S19 13.0	110 36.7	10.7	9 11.6	15.2	59.6	45	03 00	03 44	04 18	26 06	26 00	02 39	03 02	
01	198 43.0	13.6	116 36.7	10.8	8 56.4	15.3	59.6	S 50	02 28	03 21	04 01	26 11	26 00	02 41	02 58	
02	213 42.9	14.2	122 36.7	10.8	8 41.1	15.4	59.6	52	02 11	03 10	03 52	26 16	26 00	02 42	02 57	
03	228 42.8	14.8	128 36.7	10.8	8 25.7	15.4	59.7	54	01 50	02 57	03 43	26 21	26 00	02 44	02 55	
04	243 42.6	15.4	134 36.7	10.7	8 10.3	15.4	59.7	56	01 23	02 42	03 33	26 26	26 00	02 46	02 53	
05	258 42.5	16.0	140 36.7	10.7	7 54.9	15.5	59.7	S 58	00 38	02 25	03 21	26 31	26 00	02 48	02 51	
06	273 42.4	S19 16.6	146 36.7	10.8	7 39.4	15.6	59.7	S 60	00 00	02 02	03 07	26 36	26 00	02 49	02 49	
07	288 42.2	17.2	152 36.7	10.8	7 23.8	15.6	59.7									
08	303 42.1	17.8	158 36.7	10.8	7 08.2	15.7	59.8									
S 09	318 42.0	18.3	164 36.7	10.8	6 52.5	15.7	59.8									
U 10	333 41.8	18.9	170 36.7	10.8	6 36.8	15.7	59.8									
N 11	348 41.7	19.5	176 36.7	10.8	6 21.1	15.8	59.8									
D 12	3 41.6	S19 20.1	182 36.7	10.9	6 05.3	15.9	59.8	N 72	14 40	16 19	15 47	15 10	14 38	14 04		
A 13	18 41.4	20.7	188 36.7	10.9	5 49.4	15.9	59.9	N 70	13 18	15 06	16 30	15 03	14 39	14 15		
Y 14	33 41.3	21.3	194 36.7	10.9	5 33.5	15.9	59.9	68	14 02	15 25	16 40	15 15	14 57	14 41	14 24	
15	48 41.2	21.9	200 36.7	10.9	5 17.6	16.0	59.9	66	14 30	15 41	16 48	15 04	14 52	14 42	14 31	
16	63 41.0	22.5	206 36.7	10.9	5 01.6	16.0	59.9	64	14 52	15 54	16 55	15 04	14 48	14 43	14 37	
17	78 40.9	23.1	212 36.7	10.9	4 45.6	16.0	59.9	62	15 10	16 05	17 01	14 45	14 45	14 44	14 43	
18	93 40.7	S19 23.6	218 36.7	10.9	4 29.6	16.1	59.9	60	15 24	16 14	17 06	14 38	14 42	14 45	14 48	
19	108 40.6	24.2	224 36.7	10.9	4 13.5	16.1	60.0	N 58	15 36	16 22	17 11	14 32	14 39	14 45	14 52	
20	123 40.5	24.8	230 36.7	10.9	3 57.4	16.2	60.0	56	15 47	16 30	17 16	14 26	14 36	14 46	14 56	
21	138 40.3	25.4	236 36.7	10.9	3 41.2	16.2	60.0	54	15 56	16 36	17 20	14 21	14 34	14 46	14 59	
22	153 40.2	26.0	242 36.7	10.8	3 25.0	16.2	60.0	52	16 05	16 43	17 24	14 16	14 32	14 47	15 02	
23	168 40.1	26.6	248 36.7	10.8	3 08.8	16.2	60.0	50	16 12	16 48	17 28	14 12	14 30	14 47	15 05	
19 00	183 39.9	S19 27.2	254 36.7	10.9	2 52.6	16.2	60.0	45	16 28	17 00	17 36	14 02	14 26	14 48	15 12	
01	198 39.8	27.7	260 36.7	10.9	2 36.4	16.3	60.1	N 40	16 41	17 11	17 44	13 54	14 22	14 49	15 17	
02	213 39.6	28.3	266 36.7	10.9	2 20.1	16.3	60.1	35	16 53	17 20	17 50	13 48	14 19	14 50	15 21	
03	228 39.5	28.9	272 36.7	10.8	2 03.8	16.3	60.1	30	17 03	17 28	17 57	13 42	14 17	14 51	15 25	
04	243 39.3	29.5	278 36.7	10.8	1 47.5	16.4	60.1	20	17 20	17 43	18 10	13 31	14 12	14 52	15 32	
05	258 39.2	30.1	284 36.7	10.8	1 31.1	16.3	60.1	N 10	17 35	17 57	18 23	13 22	14 08	14 53	15 39	
06	273 39.1	S19 30.6	290 36.7	10.8	1 14.8	16.4	60.1	0	17 49	18 11	18 36	13 13	14 03	14 53	15 44	
07	288 38.9	31.2	296 36.7	10.8	0 98.2	16.4	60.2	S 10	18 03	18 26	18 52	13 04	13 59	14 54	15 50	
08	303 38.8	31.8	302 36.7	10.7	0 82.0	16.4	60.2	20	18 19	18 42	19 10	12 55	13 55	14 55	15 56	
09	318 38.6	32.4	308 36.7	10.7	0 65.7	16.4	60.2	30	18 36	19 03	19 34	12 44	13 50	14 56	16 04	
M 10	333 38.5	32.9	314 36.7	10.7	0 49.4	16.4	60.2	35	18 47	19 15	19 50	12 38	13 47	14 57	16 08	
O 11	348 38.3	33.5	320 36.7	10.7	0 33.2	16.5	60.2	40	18 59	19 30	20 08	12 30	13 44	14 57	16 12	
D 12	3 38.2	S19 34.1	326 36.7	10.7	0 16.9	16.4	60.2	45	19 13	19 48	20 32	12 22	13 40	14 58	16 18	
A 13	18 38.1	34.7	332 36.7	10.7	0 00.6	16.4	60.2	S 50	19 31	20 11	21 04	12 11	13 35	14 59	16 24	
Y 14	33 37.9	35.2	338 36.7	10.7	0 14.3	16.5	60.2	52	19 39	20 22	21 22	12 07	13 33	14 59	16 27	
15	48 37.8	35.8	344 36.7	10.6	0 08.0	16.5	60.3	54	19 48	20 35	21 43	12 01	13 30	15 00	16 31	
16	63 37.6	36.4	350 36.7	10.6	0 01.7	16.5	60.3	56	19 59	20 50	22 12	11 55	13 28	15 00	16 35	
17	78 37.5	37.0	356 36.7	10.6	0 15.4	16.4	60.3	S 58	20 11	21 08	22 01	11 49	13 25	15 01	16 39	
18	93 37.3	S19 37.5	362 36.7	10.6	0 09.1	16.4	60.3	S 60	20 25	21 31	22 11	11 41	13 21	15 02	16 43	
19	108 37.2	38.1	368 36.7	10.5	0 02.8	16.5	60.3									
20	123 37.0	38.7	374 36.7	10.5	0 16.5	16.4	60.3									
21	138 36.9	39.2	380 36.7	10.5	0 30.2	16.4	60.3									
22	153 36.7	39.8	386 36.7	10.5	0 43.9	16.4	60.3									

Azimuth of the Sun, Continued

42 ^m										INCREMENTS AND CORRECTIONS										43 ^m												
42 ^m	SUN PLANETS		ARIES	MOON	v or d		v or d		v or d		43 ^m	SUN PLANETS		ARIES	MOON	v or d		v or d		v or d		43 ^m	SUN PLANETS		ARIES	MOON	v or d		v or d		v or d	
00	10 300		10 317	10 013	0-0	0-0	6-0	4-3	12-0	8-5	00	10 450		10 468	10 156	0-0	0-0	6-0	4-4	12-0	8-7	00	10 450		10 468	10 156	0-0	0-0	6-0	4-4	12-0	8-7
01	10 303		10 320	10 015	0-1	0-1	6-1	4-3	12-1	8-6	01	10 453		10 470	10 159	0-1	0-1	6-1	4-4	12-1	8-8	01	10 453		10 470	10 159	0-1	0-1	6-1	4-4	12-1	8-8
02	10 305		10 322	10 018	0-2	0-1	6-2	4-4	12-2	8-6	02	10 455		10 473	10 161	0-2	0-1	6-2	4-5	12-2	8-8	02	10 455		10 473	10 161	0-2	0-1	6-2	4-5	12-2	8-8
03	10 308		10 325	10 020	0-3	0-2	6-3	4-5	12-3	8-7	03	10 458		10 475	10 163	0-3	0-2	6-3	4-6	12-3	8-9	03	10 458		10 475	10 163	0-3	0-2	6-3	4-6	12-3	8-9
04	10 310		10 327	10 023	0-4	0-3	6-4	4-5	12-4	8-8	04	10 460		10 478	10 166	0-4	0-3	6-4	4-6	12-4	9-0	04	10 460		10 478	10 166	0-4	0-3	6-4	4-6	12-4	9-0
05	10 313		10 330	10 025	0-5	0-4	6-5	4-6	12-5	8-9	05	10 463		10 480	10 168	0-5	0-4	6-5	4-7	12-5	9-1	05	10 463		10 480	10 168	0-5	0-4	6-5	4-7	12-5	9-1
06	10 315		10 332	10 027	0-6	0-4	6-6	4-7	12-6	8-9	06	10 465		10 483	10 170	0-6	0-4	6-6	4-8	12-6	9-1	06	10 465		10 483	10 170	0-6	0-4	6-6	4-8	12-6	9-1
07	10 318		10 335	10 030	0-7	0-5	6-7	4-7	12-7	9-0	07	10 468		10 485	10 173	0-7	0-5	6-7	4-9	12-7	9-2	07	10 468		10 485	10 173	0-7	0-5	6-7	4-9	12-7	9-2
08	10 320		10 337	10 032	0-8	0-6	6-8	4-8	12-8	9-1	08	10 470		10 488	10 175	0-8	0-6	6-8	4-9	12-8	9-3	08	10 470		10 488	10 175	0-8	0-6	6-8	4-9	12-8	9-3
09	10 323		10 340	10 034	0-9	0-6	6-9	4-9	12-9	9-1	09	10 473		10 490	10 178	0-9	0-7	6-9	5-0	12-9	9-4	09	10 473		10 490	10 178	0-9	0-7	6-9	5-0	12-9	9-4
10	10 325		10 342	10 037	1-0	0-7	7-0	5-0	13-0	9-2	10	10 475		10 493	10 180	1-0	0-7	7-0	5-1	13-0	9-4	10	10 475		10 493	10 180	1-0	0-7	7-0	5-1	13-0	9-4
11	10 328		10 345	10 039	1-1	0-8	7-1	5-0	13-1	9-3	11	10 478		10 495	10 182	1-1	0-8	7-1	5-1	13-1	9-5	11	10 478		10 495	10 182	1-1	0-8	7-1	5-1	13-1	9-5
12	10 330		10 347	10 042	1-2	0-9	7-2	5-1	13-2	9-4	12	10 480		10 498	10 185	1-2	0-9	7-2	5-2	13-2	9-6	12	10 480		10 498	10 185	1-2	0-9	7-2	5-2	13-2	9-6
13	10 333		10 350	10 044	1-3	0-9	7-3	5-2	13-3	9-4	13	10 483		10 500	10 187	1-3	0-9	7-3	5-3	13-3	9-6	13	10 483		10 500	10 187	1-3	0-9	7-3	5-3	13-3	9-6
14	10 335		10 352	10 046	1-4	1-0	7-4	5-2	13-4	9-5	14	10 485		10 503	10 190	1-4	1-0	7-4	5-4	13-4	9-7	14	10 485		10 503	10 190	1-4	1-0	7-4	5-4	13-4	9-7
15	10 338		10 355	10 049	1-5	1-1	7-5	5-3	13-5	9-6	15	10 488		10 505	10 192	1-5	1-1	7-5	5-4	13-5	9-8	15	10 488		10 505	10 192	1-5	1-1	7-5	5-4	13-5	9-8
16	10 340		10 357	10 051	1-6	1-1	7-6	5-4	13-6	9-6	16	10 490		10 508	10 194	1-6	1-2	7-6	5-5	13-6	9-9	16	10 490		10 508	10 194	1-6	1-2	7-6	5-5	13-6	9-9
17	10 343		10 360	10 054	1-7	1-2	7-7	5-5	13-7	9-7	17	10 493		10 510	10 197	1-7	1-2	7-7	5-6	13-7	9-9	17	10 493		10 510	10 197	1-7	1-2	7-7	5-6	13-7	9-9
18	10 345		10 362	10 056	1-8	1-3	7-8	5-5	13-8	9-8	18	10 495		10 513	10 199	1-8	1-3	7-8	5-7	13-8	10-0	18	10 495		10 513	10 199	1-8	1-3	7-8	5-7	13-8	10-0
19	10 348		10 365	10 058	1-9	1-3	7-9	5-6	13-9	9-8	19	10 498		10 515	10 202	1-9	1-4	7-9	5-7	13-9	10-1	19	10 498		10 515	10 202	1-9	1-4	7-9	5-7	13-9	10-1
20	10 350		10 367	10 061	2-0	1-4	8-0	5-7	14-0	9-9	20	10 500		10 518	10 204	2-0	1-5	8-0	5-8	14-0	10-2	20	10 500		10 518	10 204	2-0	1-5	8-0	5-8	14-0	10-2
21	10 353		10 370	10 063	2-1	1-5	8-1	5-7	14-1	10-0	21	10 503		10 520	10 206	2-1	1-5	8-1	5-9	14-1	10-2	21	10 503		10 520	10 206	2-1	1-5	8-1	5-9	14-1	10-2
22	10 355		10 372	10 065	2-2	1-6	8-2	5-8	14-2	10-1	22	10 505		10 523	10 209	2-2	1-6	8-2	5-9	14-2	10-3	22	10 505		10 523	10 209	2-2	1-6	8-2	5-9	14-2	10-3
23	10 358		10 375	10 068	2-3	1-6	8-3	5-9	14-3	10-1	23	10 508		10 525	10 211	2-3	1-7	8-3	6-0	14-3	10-4	23	10 508		10 525	10 211	2-3	1-7	8-3	6-0	14-3	10-4
24	10 360		10 377	10 070	2-4	1-7	8-4	6-0	14-4	10-2	24	10 510		10 528	10 213	2-4	1-7	8-4	6-1	14-4	10-4	24	10 510		10 528	10 213	2-4	1-7	8-4	6-1	14-4	10-4
25	10 363		10 380	10 073	2-5	1-8	8-5	6-0	14-5	10-3	25	10 513		10 530	10 216	2-5	1-8	8-5	6-2	14-5	10-5	25	10 513		10 530	10 216	2-5	1-8	8-5	6-2	14-5	10-5
26	10 365		10 382	10 075	2-6	1-8	8-6	6-1	14-6	10-3	26	10 515		10 533	10 218	2-6	1-9	8-6	6-2	14-6	10-6	26	10 515		10 533	10 218	2-6	1-9	8-6	6-2	14-6	10-6
27	10 368		10 385	10 077	2-7	1-9	8-7	6-2	14-7	10-4	27	10 518		10 535	10 221	2-7	2-0	8-7	6-3	14-7	10-7	27	10 518		10 535	10 221	2-7	2-0	8-7	6-3	14-7	10-7
28	10 370		10 387	10 080	2-8	2-0	8-8	6-2	14-8	10-5	28	10 520		10 538	10 223	2-8	2-0	8-8	6-4	14-8	10-7	28	10 520		10 538	10 223	2-8	2-0	8-8	6-4	14-8	10-7
29	10 373		10 390	10 082	2-9	2-1	8-9	6-3	14-9	10-6	29	10 523		10 540	10 225	2-9	2-1	8-9	6-5	14-9	10-8	29	10 523		10 540	10 225	2-9	2-1	8-9	6-5	14-9	10-8
30	10 375		10 392	10 085	3-0	2-1	9-0	6-4	15-0	10-6	30	10 525		10 543	10 228	3-0	2-2	9-0	6-5	15-0	10-9	30	10 525		10 543	10 228	3-0	2-2	9-0	6-5	15-0	10-9
31	10 378		10 395	10 087	3-1	2-2	9-1	6-4	15-1	10-7	31	10 528		10 545	10 230	3-1	2-2	9-1	6-6	15-1	10-9	31	10 528		10 545	10 230	3-1	2-2	9-1	6-6	15-1	10-9
32	10 380		10 397	10 089	3-2	2-3	9-2	6-5	15-2	10-8	32	10 530		10 548	10 233	3-2	2-3	9-2	6-7	15-2	11-0	32	10 530		10 548	10 233	3-2	2-3	9-2	6-7	15-2	11-0
33	10 383		10 400	10 092	3-3	2-3	9-3	6-6	15-3	10-8	33	10 533		10 550	10 235	3-3	2-4	9-3	6-7	15-3	11-1	33	10 533		10 550	10 235	3-3	2-4	9-3	6-7	15-3	11-1
34	10 385		10 402	10 094	3-4	2-4	9-4	6-7	15-4	10-9	34	10 535		10 553	10 237	3-4	2-5	9-4	6-8	15-4	11-2	34	10 535		10 553	10 237	3-4	2-5	9-4	6-8	15-4	11-2
35	10 388		10 405	10 097	3-5	2-5	9-5	6-7	15-5	11-0	35	10 538		10 555	10 240	3-5	2-5	9-5	6-9	15-5	11-2	35	10 538		10 555	10 240	3-5	2-5	9-5	6-9	15-5	11-2
36	10 390		10 407	10 099	3-6	2-6	9-6	6-8	15-6	11-1	36	10 540		10 558	10 242	3-6	2-6	9-6	7-0	15-6	11-3	36	10 540		10 558	10 242	3-6	2-6	9-6	7-0	15-6	11-3
37	10 393		10 410	10 101	3-7	2-6	9-7	6-9	15-7	11-1	37	10 543		10 560	10 244	3-7	2-7	9-7	7-0	15-7	11-4	37	10 543		10 560	10 244	3-7	2-7	9-7	7-0	15-7	11-4
38	10 395		10 413	10 104	3-8	2-7	9-8																									

Azimuth of the Sun, Continued

Look at the left-hand column of the stip form below. Notice that you'll find values for Dec. Inc/Z Diff, Lat Inc/Z Diff, and LHA Inc/Z Diff. This is where we enter the leftover values from our whole degrees of DR Lat, Declination, and LHA. To do this, we must convert our leftover values into tenths of degrees by dividing each by 60 and rounding to the closest tenth of a degree. Finding Z Diff is a matter of inspecting Pub 229 (see figs. 9-3 and 9-4) in the following manner:

For Dec Inc/Z Diff note the values for the whole degree of dec that you entered the table with and the next high dec, then find the difference. Here are the values for our example problem: Dec 19° Z = 129.1 Dec 20° Z = 129.8. The difference between the values is 0.7. Since the value is increasing between 19° and 20° , we assign it a positive value (+).

Repeat the same procedure for finding Z Diff for Latitude and LHA. In other words, compare lat. 33° and lat 34° . Then compare LHA 51° and 52° .

OPNAV 3130/ Azimuth by 229	ACTION	Completed Strip Form
Tab Z	Enter Pub 229 with entering arguments of Lat 33° , LHA 51° , and Dee 19° . Make sure that you enter on the portion of the page that indicates LATITUDE CONTRARY TO DECLINATION . Follow 19° of Dee across the page to where it falls under the 33° Latitude column and record the value for Tab Z.	129.1
Dec Inc/Z Diff	Dec. Inc (left) = $31.6' + 60$ which = .53 rounded to 0.5. Compares Z's for Z Diff.	0.5 / + 0.7
Dec Corr	Multiply Dee Inc by Z Diff.	+ 0.35
Lat Inc/Z Diff	Lat Inc (left) = $37.0' + 60$ which = .62 rounded to 0.6. Compares Z's for Z Diff.	0.6 / + 0.3
Lat Corr	Multiply Lat Inc by Z Diff.	+ 0.18
LHA Inc/Z Diff	LHA Inc (left) = $53.4' + 60$ which = .89 rounded to 0.9. Compares Z's for Z Diff.	0.9 / - 0.7
LHA Corr	Multiply LHA Inc by 2 Diff.	- 0.63
Dec Corr	Drop the Dec Corr Down.	+ 0.35
Lat Corr	Drop the Lat Corr Down.	+ 0.18
Total corr	Add the LHA, Dec, and Lat Corr.	- 0.10

We have now accounted for our leftover values and now can find the Exact Z.

Azimuth of the Sun, Continued

LATITUDE CONTRARY NAME TO DECLINATION																								L.H.A. 51°, 309°	
Dec.	30°			31°			32°			33°			34°			35°			36°			37°			Dec.
0	Hc	d	Z	Hc	d	Z	Hc	d	Z	Hc	d	Z	Hc	d	Z	Hc	d	Z	Hc	d	Z	Hc	d	Z	0
0	33 01.5	36.0	112.0	32 38.7	36.9	112.6	32 15.3	37.8	113.2	31 51.4	38.7	113.8	31 26.9	39.5	114.4	31 01.9	40.3	114.9	30 36.4	41.2	115.5	30 10.3	41.9	116.0	0
1	32 25.5	36.1	113.0	32 01.8	37.0	113.6	31 37.5	38.2	114.1	31 12.7	39.0	114.7	30 47.4	39.9	115.2	30 21.6	40.7	115.8	29 55.2	41.4	116.3	29 28.4	42.2	116.8	1
2	31 49.1	36.2	113.9	31 24.5	37.1	114.5	30 59.3	38.5	115.0	30 33.7	39.4	115.6	30 07.5	40.2	116.1	29 40.9	41.0	116.6	29 13.8	41.8	117.1	28 46.2	42.5	117.6	2
3	31 12.3	37.3	114.9	30 46.8	38.1	115.4	30 20.8	38.9	115.9	29 54.3	39.7	116.5	29 27.3	40.5	117.0	28 59.9	41.3	117.5	28 32.0	42.0	117.9	28 03.7	42.8	118.4	3
4	30 35.0	37.6	115.8	30 08.7	38.4	116.3	29 41.9	39.3	116.8	29 14.6	40.1	117.3	28 46.8	40.8	117.8	28 18.6	41.6	118.3	27 50.0	42.4	118.8	27 20.9	43.1	119.2	4
5	29 57.4	37.9	116.7	29 30.3	38.8	117.2	29 02.6	39.6	117.7	28 34.5	40.3	118.2	28 06.0	41.1	118.6	27 37.0	41.9	119.1	27 07.6	42.6	119.6	26 37.8	43.3	120.0	5
6	29 19.5	38.3	117.6	28 51.5	39.1	118.1	28 23.0	39.9	118.5	27 54.2	40.7	119.0	27 24.9	41.5	119.5	26 55.1	42.1	119.9	26 25.0	42.8	120.3	25 54.5	43.5	120.8	6
7	28 41.2	38.7	118.4	28 12.4	39.5	118.9	27 43.1	40.2	119.4	27 13.5	41.0	119.8	26 43.4	41.7	120.3	26 13.0	42.4	120.7	25 42.2	43.1	121.1	25 11.0	43.8	121.5	7
8	28 02.5	39.0	119.3	27 32.9	39.7	119.8	27 02.9	40.5	120.2	26 32.5	41.2	120.7	26 01.7	41.9	121.1	25 30.6	42.7	121.5	24 59.1	43.4	121.9	24 27.2	44.0	122.3	8
9	27 23.5	39.3	120.2	26 53.7	40.1	120.6	26 22.4	40.8	121.0	25 51.3	41.5	121.5	25 19.8	42.2	121.9	24 47.9	42.9	122.3	24 15.7	43.5	122.7	23 43.2	44.2	123.0	9
10	26 44.2	39.5	121.0	26 13.1	40.3	121.4	25 41.6	41.0	121.9	25 09.8	41.8	122.3	24 37.6	42.5	122.7	24 05.0	43.1	123.0	23 32.2	43.8	123.4	22 59.0	44.4	123.8	10
11	26 04.7	39.9	121.9	25 32.8	40.6	122.3	25 00.6	41.3	122.7	24 28.0	42.0	123.1	23 55.1	42.8	123.4	23 21.9	43.3	123.8	22 48.4	44.0	124.2	22 14.6	44.6	124.5	11
12	25 24.8	40.2	122.7	24 52.2	40.9	123.1	24 19.3	41.6	123.5	23 46.0	42.2	123.8	23 12.5	42.9	124.2	22 38.6	43.6	124.5	22 04.4	44.2	124.9	21 30.0	44.8	125.2	12
13	24 44.6	40.4	123.5	24 11.3	41.1	123.9	23 37.7	41.8	124.3	23 03.8	42.5	124.6	22 29.6	43.1	125.0	21 55.0	43.7	125.3	21 20.2	44.3	125.6	20 45.2	45.0	125.9	13
14	24 04.2	40.7	124.3	23 30.2	41.3	124.7	22 55.9	42.0	125.0	22 21.3	42.6	125.4	21 46.5	43.4	125.7	21 11.3	43.9	126.0	20 35.9	44.6	126.3	20 00.2	45.1	126.6	14
15	23 23.5	40.9	125.1	22 48.9	41.6	125.5	22 13.9	42.2	125.8	21 38.7	42.9	126.1	21 03.1	43.5	126.5	20 27.4	44.1	126.8	19 51.3	44.7	127.0	19 15.1	45.3	127.3	15
16	22 42.6	41.1	125.9	22 07.3	41.8	126.3	21 31.7	42.5	126.6	20 55.8	43.1	126.9	20 19.6	43.6	127.2	19 43.3	44.3	127.5	19 06.6	44.8	127.8	18 29.8	45.5	128.0	16
17	22 01.5	41.3	126.7	21 25.5	42.0	127.0	20 49.2	42.6	127.3	20 12.7	43.2	127.6	19 36.0	43.9	127.9	18 59.0	44.5	128.2	18 21.8	45.1	128.5	17 44.3	45.8	128.7	17
18	21 20.1	41.6	127.5	20 43.5	42.2	127.8	20 06.6	42.8	128.1	19 29.5	43.5	128.4	18 52.1	44.0	128.6	18 14.5	44.6	128.9	17 36.7	45.1	129.2	16 58.8	45.9	129.4	18
19	20 38.5	41.8	128.3	20 01.3	42.5	128.5	19 23.8	43.1	128.8	18 46.0	43.6	129.1	18 08.1	44.2	129.4	17 29.9	44.7	129.6	16 51.6	45.3	129.8	16 13.0	45.8	130.1	19
20	19 56.7	41.9	129.0	19 18.8	42.5	129.3	18 40.7	43.1	129.6	18 02.4	43.7	129.8	17 23.9	44.3	130.1	16 45.2	44.9	130.3	16 06.3	45.5	130.5	15 27.2	46.0	130.7	20
21	19 14.8	42.2	129.8	18 36.3	42.8	130.0	17 57.6	43.4	130.3	17 18.7	44.0	130.5	16 39.6	44.5	130.8	16 00.3	45.0	131.0	15 20.8	45.5	131.2	14 41.2	46.0	131.4	21
22	18 32.6	42.4	130.5	17 53.5	42.9	130.8	17 14.2	43.5	131.0	16 34.7	44.0	131.3	15 55.1	44.8	131.5	15 15.3	45.3	131.7	14 35.3	45.7	131.9	13 55.2	46.2	132.1	22
23	17 50.2	42.5	131.3	17 10.6	43.1	131.5	16 30.7	43.6	131.7	15 50.7	44.2	132.0	15 10.5	44.8	132.2	14 30.1	45.2	132.4	13 49.6	45.8	132.5	13 09.0	46.6	132.7	23
24	17 07.7	42.6	132.0	16 27.5	43.2	132.2	15 47.1	43.8	132.4	15 06.1	44.3	132.7	14 25.7	44.9	132.9	13 44.1	45.4	133.1	13 02.4	46.0	133.2	12 20.7	46.8	133.4	24
25	16 25.1	42.7	132.7	15 44.2	43.3	132.9	15 03.4	44.0	133.0	14 21.6	44.5	133.2	13 40.1	45.0	133.4	13 00.0	45.5	133.6	12 19.1	46.1	133.7	11 37.5	47.0	133.9	25
26	15 42.3	42.8	133.4	15 00.9	43.4	133.6	14 19.5	44.1	133.7	13 37.6	44.6	133.9	12 55.9	45.1	134.1	12 14.1	45.6	134.3	11 31.9	46.2	134.5	10 49.4	47.5	134.7	26
27	15 00.0	42.9	134.1	14 17.6	43.5	134.8	13 36.2	44.2	134.9	12 54.3	44.7	135.0	12 12.8	45.2	135.2	11 30.6	45.7	135.4	10 47.8	46.3	135.6	10 04.4	48.0	135.8	27
28	14 17.6	43.0	134.8	13 34.2	43.6	135.5	12 51.7	44.3	135.6	12 10.0	44.8	135.7	11 28.3	45.3	135.9	10 46.1	45.8	136.1	10 03.3	46.4	136.3	9 20.0	48.5	136.5	28
29	13 35.0	43.1	135.5	12 51.5	43.7	136.2	12 08.4	44.4	136.3	11 26.5	44.9	136.4	10 44.4	45.4	136.6	10 02.2	45.9	136.8	9 19.9	46.5	137.0	8 37.5	49.0	137.2	29
30	12 52.3	43.2	136.2	12 08.2	43.8	136.9	11 24.9	44.5	137.0	10 42.8	45.0	137.1	10 00.6	45.5	137.3	9 18.3	46.0	137.5	8 35.9	46.6	137.7	7 53.4	49.5	137.9	30
31	12 09.5	43.3	136.9	11 24.9	43.9	137.6	10 42.5	44.6	137.7	10 00.3	45.1	137.8	9 18.0	45.6	138.0	8 35.6	46.1	138.2	7 53.2	46.7	138.4	7 10.7	50.0	138.6	31
32	11 26.7	43.4	137.6	10 41.0	44.0	138.3	9 58.6	44.7	138.4	9 16.4	45.2	138.5	8 34.1	45.7	138.7	7 51.6	46.2	138.9	7 09.2	46.8	139.1	6 26.7	50.5	139.3	32
33	10 43.8	43.5	138.3	9 58.1	44.1	139.0	9 15.7	44.8	139.1	8 33.5	45.3	139.2	7 51.0	45.8	139.4	7 08.5	46.3	139.6	6 26.0	46.9	139.8	5 43.5	51.0	139.9	33
34	10 00.9	43.6	139.0	9 14.2	44.2	139.7	8 31.8	44.9	139.8	7 49.6	45.4	139.9	7 06.9	45.9	140.1	6 24.4	46.4	140.3	5 41.9	47.0	140.5	5 00.0	51.5	140.6	34
35	9 17.9	43.7	139.7	8 31.2	44.3	140.4	7 48.8	45.0	140.5	7 06.6	45.5	140.6	6 23.9	46.0	140.8	5 41.4	46.5	141.0	4 58.9	47.1	141.2	4 16.4	52.0	141.3	35
36	8 34.8	43.8	140.4	7 48.2	44.4	141.1	7 05.3	45.1	141.2	6 23.1	45.6	141.3	5 40.4	46.1	141.5	4 57.9	46.6	141.7	4 15.4	47.2	141.9	3 32.9	52.5	142.0	36
37	7 51.8	43.9	141.1	7 04.7	44.5	141.8	6 22.8	45.2	141.9	5 40.6	45.7	142.0	4 57.7	46.2	142.2	4 15.2	46.7	142.4	3 32.7	47.3	142.6	2 50.2	53.0	142.7	37
38	7 08.7	44.0	141.8	6 21.6	44.6	142.5	5 39.7	45.3	142.6	4 57.5	45.8	142.7	4 14.6	46.3	142.9	3 32.1	46.8	143.1	2 49.6	47.4	143.3	2 07.1	53.5	143.4	38
39	6 25.7	44.1	142.5	5 38.6	44.7	143.2	4 56.6	45.4	143.3	4 14.4	45.9	143.4	3 31.5	46.4	143.6	2 49.0	46.9	143.8	2 06.5	47.5	144.0	1 24.0	54.0	144.1	39
40	5 42.6	44.2	143.2	4 55.5	44.8	143.9	4 13.5	45.5	144.0	3 31.3	46.0	144.1	2 48.4	46.5	144.3	2 05.9	47.0	144.5	1 23.4	47.6	144.7	0 40.9	54.5	144.8	40
41	4 59.5	44.3	143.9	4 12.4	44.9	144.6	3 30.4	45.6	144.7	2 48.2	46.1	4													

Azimuth of the Sun, Continued

LATITUDE CONTRARY NAME TO DECLINATION																								L.H.A. 52°, 308°									
Dec.	30°			31°			32°			33°			34°			35°			36°			37°			Dec.								
	M.	d	Z	M.	d	Z	M.	d	Z	M.	d	Z	M.	d	Z	M.	d	Z	M.	d	Z	M.	d	Z									
0	32	13.2	356	111.3	31	51.1	366	111.9	31	28.4	374	112.5	31	05.2	383	113.1	30	41.5	392	113.6	30	17.2	400	114.1	29	52.4	408	114.7	29	27.1	416	115.2	0
1	31	37.6	366	112.3	31	14.5	376	112.8	30	51.0	374	113.4	30	26.9	383	113.9	29	02.3	392	114.5	29	37.2	400	115.0	29	11.6	408	115.5	28	45.5	416	116.0	1
2	31	01.5	366	113.2	30	37.5	376	113.8	30	13.1	384	114.3	29	48.2	392	114.8	29	22.7	398	115.3	28	56.8	406	115.8	28	30.4	414	116.3	28	03.6	422	116.8	2
3	30	25.0	366	114.1	30	00.2	377	114.7	29	34.9	384	115.2	29	09.1	394	115.7	28	42.9	402	116.2	28	16.1	409	116.7	27	49.0	416	117.2	27	21.4	422	117.7	3
4	29	48.1	372	115.1	29	22.5	384	115.6	28	56.3	390	116.1	28	29.7	397	116.6	28	02.7	403	117.0	27	35.2	411	117.5	27	07.2	420	118.0	26	38.9	427	118.4	4
5	29	10.9	376	116.0	28	44.4	385	116.5	28	17.4	394	116.9	27	50.0	400	117.4	27	22.2	406	117.9	26	53.9	413	118.3	26	25.2	421	118.8	25	56.2	430	119.2	5
6	28	33.3	380	116.8	28	35.9	387	117.3	27	38.2	396	117.8	27	10.0	403	118.3	26	41.4	411	118.7	26	12.4	418	119.1	25	43.0	426	119.6	25	13.2	435	120.0	6
7	27	55.3	383	117.7	27	27.2	391	118.2	26	58.6	398	118.6	26	29.7	406	119.1	26	00.3	413	119.5	25	30.6	421	119.9	25	00.4	427	120.3	24	30.0	435	120.7	7
8	27	17.0	385	118.6	26	48.1	393	119.0	26	18.8	401	119.5	25	49.1	408	119.9	25	19.0	416	120.3	24	48.5	423	120.7	24	17.7	430	121.1	23	46.5	437	121.5	8
9	26	38.3	388	119.5	26	08.8	397	119.9	25	38.7	404	120.3	25	08.2	411	120.7	24	37.4	419	121.1	24	06.2	425	121.5	23	34.7	432	121.9	23	02.8	441	122.2	9
10	25	59.6	392	120.3	25	29.1	400	120.7	24	58.3	407	121.1	24	27.1	414	121.5	23	55.5	420	121.9	23	23.7	428	122.3	22	51.5	435	122.6	22	19.0	444	123.0	10
11	25	20.4	395	121.1	24	49.1	403	121.5	24	17.6	410	121.9	23	45.7	417	122.3	23	13.5	424	122.7	22	40.9	430	123.0	22	08.0	438	123.4	21	34.9	443	123.7	11
12	24	40.9	398	122.0	24	08.9	406	122.4	23	36.7	413	122.7	23	04.1	420	123.1	22	31.1	427	123.4	21	57.9	434	123.8	21	24.4	441	124.1	20	50.6	446	124.4	12
13	24	01.1	400	122.8	23	28.5	409	123.2	22	55.5	416	123.5	22	22.2	423	123.9	21	48.6	430	124.2	21	14.7	437	124.5	20	40.6	444	124.8	20	06.2	452	125.2	13
14	23	21.1	402	123.6	22	47.8	412	124.0	22	14.1	419	124.3	21	40.1	426	124.6	20	55.9	433	125.0	20	31.4	440	125.3	19	56.6	446	125.6	19	21.6	458	126.3	14
15	22	40.9	405	124.4	22	06.8	415	124.8	21	32.5	421	125.1	20	57.8	428	125.4	20	22.9	435	125.7	19	47.8	442	126.0	19	12.4	449	126.3	18	36.8	466	126.9	15
16	22	00.4	408	125.2	21	25.6	418	125.5	20	50.6	424	125.9	20	15.3	432	126.2	19	39.8	439	126.4	19	04.1	446	126.7	18	28.1	453	127.0	17	51.8	473	127.3	16
17	21	19.6	409	126.0	20	44.2	420	126.3	20	08.6	426	126.6	19	32.7	439	126.9	18	56.5	443	127.2	18	20.2	449	127.4	17	43.6	457	127.7	17	06.8	481	128.0	17
18	20	38.7	412	126.8	20	02.6	422	127.1	19	26.3	428	127.4	18	49.8	435	127.6	18	13.1	442	127.9	17	36.1	449	128.2	16	58.9	461	128.4	16	21.5	493	128.6	18
19	19	57.5	413	127.6	19	20.8	423	127.8	18	43.9	429	128.1	18	06.8	437	128.4	17	29.4	444	128.6	16	51.9	451	128.9	16	14.1	464	129.1	15	36.5	505	129.3	19
20	19	16.2	416	128.3	18	38.9	425	128.6	18	01.3	430	128.9	17	23.6	439	129.1	16	45.6	447	129.3	16	07.5	455	129.6	15	29.2	465	129.8	14	50.7	516	130.0	20
21	18	34.6	417	129.1	17	56.7	426	129.3	17	18.5	431	129.6	16	40.2	440	129.8	16	01.7	449	130.1	15	23.0	456	130.3	14	44.1	471	130.5	14	05.1	527	130.7	21
22	17	52.9	419	129.9	17	14.4	428	130.1	16	35.6	433	130.3	15	56.7	442	130.5	15	17.6	451	130.8	14	38.4	458	131.0	13	58.9	473	131.2	13	19.4	539	131.3	22
23	17	11.0	421	130.6	16	31.9	430	130.8	15	52.5	435	131.1	15	13.0	447	131.3	14	33.4	456	131.5	13	53.6	466	131.6	12	13.7	485	131.8	12	33.8	550	132.0	23
24	16	28.9			15	49.3																											
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Azimuth of the Sun, Continued

Tab Z	Drop the value from Tab Z.	129.1
Exact Z (-360)	Apply the Total Corr to Tab Z (- 0.10).	129.0
Exact Zn	<p>On each page of Pub 229 are small notes that state:</p> <p>In N. Lat. if LHA is Greater than 180 then Zn= Z if LHA is Less than 180 then Zn= 360-Z</p> <p>In S. Lat. if LHA is Greater than 180 then Zn= 180-Z if LHA is Less than 180 then Zn= 180+Z</p> <p>Our Lat is N. and LHA is less than 180, therefore Zn=360 - 129 or 231°.</p>	231.0
Gyro Bearing	Enter the gyro bearing from the observation.	231.6
Gyro Error	<p>Find the difference between Exact Zn and the gyro bearing and name the error. If the gyro bearing is less than the Exact Zn, the error is easterly, if more than Exact Zn, it's westerly. Use this memory aid:</p> <p>Gyro least - error east, Gyro best - error west.</p>	0.6 West

We have now used the Sun to find the error on our gyrocompass. As stated before, a greater degree of accuracy can be obtained by making several observations and then working the solutions and averaging the results. This may seem a bit tedious, however, you may work all observations at once. This is easily accomplished by entering data in the strip form in stages.

Try this method. First enter GMT DR Lat, DR Long, GHA, Dec, and d#. Next find Incements (m/s), LHA, and True Dec. Now find your leftover values for Dee Inc, Lat Inc, and LHA Inc and enter Pub 229.

Once you have completed the solutions for all observations, you can average the results. Here's an example:

Error 1 = .6 W Error 2 = .5 W Error 3 = .7 W for a total of $1.8 \div 3 = .6$ W

Azimuth by Polaris

Azimuth by Polaris

Polaris (the North Star) is always within about 2° of true north. The true azimuth of Polaris is tabulated in the *Nautical Almanac* in the Polaris Tables for northern latitudes up to 65°.

Gathering Information: The entering arguments for the Polaris Tables are the LHA of Aries (GHA of Aries plus east longitude or minus west longitude) and latitude (at intervals of 5°, 10° or 20°). An extract from the *Nautical Almanac* Polaris azimuth table, which appears at the foot of the Polaris Tables, is shown in figure 9-5. As you can see, the interpolation can be done by visual inspection of the appropriate LHA and latitude.

The normal use of Polaris for obtaining compass error is when your ship is in the lower northern latitudes. This allows you to take a bearing on Polaris using the telescopic alidade. Since the computation and interpolation of azimuth by Polaris are relatively simple, we will not go into step-by-step procedures in this text.

POLARIS (POLE STAR) TABLES, 1983												275
FOR DETERMINING LATITUDE FROM SEXTANT ALTITUDE AND FOR AZIMUTH												
L.H.A. ARIES	120°– 129°	130°– 139°	140°– 149°	150°– 159°	160°– 169°	170°– 179°	180°– 189°	190°– 199°	200°– 209°	210°– 219°	220°– 229°	230°– 239°
Lat.	AZIMUTH											
°	°	°	°	°	°	°	°	°	°	°	°	°
0	359·2	359·2	359·2	359·3	359·4	359·5	359·6	359·7	359·9	0·0	0·2	0·3
20	359·1	359·2	359·2	359·3	359·4	359·5	359·6	359·7	359·9	0·0	0·2	0·3
40	358·9	359·0	359·0	359·1	359·2	359·3	359·5	359·7	359·8	0·0	0·2	0·4
50	358·7	358·8	358·8	358·9	359·1	359·2	359·4	359·6	359·8	0·0	0·2	0·5
55	358·6	358·6	358·7	358·8	359·0	359·1	359·3	359·6	359·8	0·0	0·3	0·5
60	358·4	358·4	358·5	358·6	358·8	359·0	359·2	359·5	359·8	0·0	0·3	0·6
65	358·1	358·1	358·2	358·4	358·6	358·8	359·1	359·4	359·7	0·0	0·4	0·7

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Figure 9-5. Extract from the *Nautical Almanac* Polaris Tables.

Amplitude of the Sun

Amplitude

An amplitude of the Sun or other celestial body can be used to determine gyro error. An amplitude (A) is the arc of the horizon between the prime vertical circle (the vertical circle through the east and west points of the horizon) and the observed body. The prime vertical circle may be true or magnetic depending upon which east or west points are involved. If the body is observed when its center is on the celestial horizon, the amplitude can be taken directly from table 27 of *Bowditch*, Volume II.

Horizons

The celestial horizon differs from the one you see (the visible horizon) because it runs through the center of Earth. There are a lot of computations that must be done to determine the celestial horizon of a body, but for now we will just say that it is the horizon that a navigator uses for all celestial computations.

When the center of the Sun is on the celestial horizon, its lower limb (lower edge) is about two-thirds of the diameter of the Sun above the visible horizon. When the center of the Moon is on the celestial horizon, its upper limb (upper edge) is on the visible horizon.

Figure 9-6 shows the relationship of the visible horizon to the celestial horizon. When planets and stars are on the celestial horizon, they are a little more than one Sun diameter above the visible horizon.

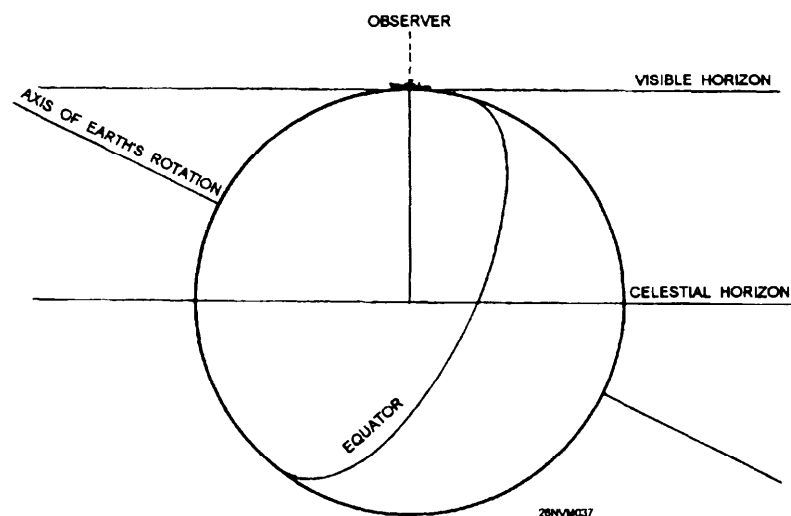


Figure 9-6. The visible and celestial horizons.

Amplitude of the Sun, Continued

Labeling the Amplitude

The amplitude of a body is given the prefix E (east) if the body is rising and the prefix W (west) if the body is setting. Additionally, the amplitude of a body is given the suffix N (north) if the body has northerly declination and the suffix S (south) if it has southerly declination.

Finding Amplitude of the Sun Using the Celestial Horizon

As discussed above, the amplitude of a body can be taken directly from table 27 of *Bowditch*, Volume II, if the body is observed when its center is on the celestial horizon. Since the Sun is most commonly used for amplitudes, it will be the topic of our discussion.

Gathering Information: To observe the Sun when it is on the celestial horizon, its lower limb must be about two-thirds of the diameter above the visible horizon. You must know the Greenwich mean time (GMT) of your observation to determine the Sun's declination from the right-hand daily pages of the *Nautical Almanac*, your DR Lat. at the time of observation, and the true bearing of the Sun as observed using a telescopic alidade.

Example Problem

The DR latitude of your ship is $51^{\circ}04.6'N$. The declination of the setting Sun was $19^{\circ}00.4'N$. Your true bearing (as observed by a telescopic alidade) to the Sun was 300° .

From this known information, we can use table 27 of *Bowditch* to determine the amplitude.

Figure 9-7 shows an excerpt from table 27. By inspection of figure 9-7, you can see that you must enter the left-hand column with your ship's DR latitude. You can also see that the Sun's declination is listed across the top of the table. Since latitude 51° and declination 19° are closest to our entering values, we determine that the amplitude of the Sun when it is on the celestial horizon is 31.2° . Now that we have the amplitude, what do we do with it? First of all, there are some basic rules that must be applied that relate to our previous discussion of the assigned prefix and suffix of an amplitude. Our amplitude was taken when the Sun was setting, and its declination name is north. Using the rules for labeling the amplitude, we label the amplitude as follows: $W\ 31.2^{\circ}\ N$. We use W because the Sun is setting and N because the Sun's declination is N.

Amplitude of the Sun, Continued

TABLE 27 Amplitudes														
Latitude	Declination													Latitude
	18°0	18°5	19°0	19°5	20°0	20°5	21°0	21°5	22°0	22°5	23°0	23°5	24°0	
0	18.0	18.5	19.0	19.5	20.0	20.5	21.0	21.5	22.0	22.5	23.0	23.5	24.0	0
10	18.3	18.8	19.3	19.8	20.3	20.8	21.3	21.8	22.4	22.9	23.4	23.9	24.4	10
15	18.7	19.2	19.7	20.2	20.7	21.3	21.8	22.3	22.8	23.3	23.9	24.4	24.9	15
20	19.2	19.7	20.3	20.8	21.3	21.9	22.4	23.0	23.5	24.0	24.6	25.1	25.6	20
25	19.9	20.5	21.1	21.6	22.2	22.7	23.3	23.9	24.4	25.0	25.5	26.1	26.7	25
30	20.9	21.5	22.1	22.7	23.3	23.9	24.4	25.0	25.6	26.2	26.8	27.4	28.0	30
32	21.4	22.0	22.6	23.2	23.8	24.4	25.0	25.6	26.2	26.8	27.4	28.0	28.7	32
34	21.9	22.5	23.1	23.7	24.4	25.0	25.6	26.2	26.9	27.5	28.1	28.7	29.4	34
36	22.5	23.1	23.7	24.4	25.0	25.7	26.3	26.9	27.6	28.2	28.9	29.5	30.2	36
38	23.1	23.7	24.4	25.1	25.7	26.4	27.1	27.7	28.4	29.1	29.7	30.4	31.1	38
40	23.8	24.5	25.2	25.8	26.5	27.2	27.9	28.6	29.3	30.0	30.7	31.4	32.1	40
41	24.2	24.9	25.6	26.3	26.9	27.6	28.3	29.1	29.8	30.5	31.2	31.9	32.6	41
42	24.6	25.3	26.0	26.7	27.4	28.1	28.8	29.5	30.3	31.0	31.7	32.5	33.2	42
43	25.0	25.7	26.4	27.2	27.9	28.6	29.3	30.1	30.8	31.6	32.3	33.0	33.8	43
44	25.4	26.2	26.9	27.6	28.4	29.1	29.9	30.6	31.4	32.1	32.9	33.7	34.4	44
45	25.9	26.7	27.4	28.2	28.9	29.7	30.5	31.2	32.0	32.8	33.5	34.3	35.1	45
46	26.4	27.2	27.9	28.7	29.5	30.3	31.1	31.8	32.6	33.4	34.2	35.0	35.8	46
47	26.9	27.7	28.5	29.3	30.1	30.9	31.7	32.5	33.3	34.1	35.0	35.8	36.6	47
48	27.5	28.3	29.1	29.9	30.7	31.6	32.4	33.2	34.0	34.9	35.7	36.6	37.4	48
49	28.1	28.9	29.8	30.6	31.4	32.3	33.1	34.0	34.8	35.7	36.6	37.4	38.3	49
50	28.7	29.6	30.4	31.3	32.1	33.0	33.9	34.8	35.6	36.5	37.4	38.3	39.3	50
51	29.4	30.3	31.2	32.0	32.9	33.8	34.7	35.6	36.5	37.5	38.4	39.3	40.3	51
52	30.1	31.0	31.9	32.8	33.7	34.7	35.6	36.5	37.5	38.4	39.4	40.4	41.3	52
53	30.9	31.8	32.8	33.7	34.6	35.6	36.5	37.5	38.5	39.5	40.5	41.5	42.5	53
54	31.7	32.7	33.6	34.6	35.6	36.6	37.6	38.6	39.6	40.6	41.7	42.7	43.8	54
70.0	64.6	68.1	72.2	77.4	90.0									70.0
70.5	67.8	71.9	77.2	90.0										70.5
71.0	71.7	77.1	90.0											71.0
71.5	76.9	90.0												71.5
72.0	90.0													72.0

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Figure 9-7. Excerpt from table 27.

Amplitude of the Sun, Continued

Finding Amplitude of the Sun Using the Celestial Horizon, continued

With the amplitude properly labeled, we can now follow another set of rules to determine the azimuth.

Rules:

1. Rising Sun with north declination, subtract the amplitude from 090°
2. Rising Sun with south declination, add the amplitude to 090°
3. Setting Sun with north declination, add the amplitude to 270°
4. Setting Sun with south declination, subtract the amplitude from 270°

By following the rules above, our amplitude can now be converted to an azimuth as follows:

$$W31.2^\circ N + 270^\circ = 301.2^\circ$$

Our true bearing to the Sun was 300° . Gyro error can be determined as follows:

$$\begin{array}{r} 301.2^\circ \text{ (azimuth)} \\ \underline{300.0^\circ \text{ (gyro bearing)}} \\ 1.2^\circ \text{ E (gyro error)} \end{array}$$

We find the name of the error by using our memory aid

Gyro least - error EAST, Gyro best - error WEST.

Finding Amplitude Using the Visible Horizon

If the body is observed when its center is on the visible horizon, a correction from table 28 of *Bowditch*, Volume II, is applied to the value taken from *Bowditch's* table 27. Refer to table 28 for step-by-step instructions.

The Celestial LOP

General Information

You have seen how lines of position, obtained through bearings on terrestrial objects, are used to fix a ship's position in piloting. You know that a line of position (LOP) is a locus of possible positions of the ship. In other words, the ship's position must be somewhere along that line. A fix, by definition, is a relatively accurate determination of latitude and longitude. In practice, this position is the intersection of two or more lines of position; but often it is not the ship's exact position because you can always assume some errors in observation, plotting, and the like.

The celestial navigator must establish lines of position by applying the results of observations of heavenly bodies. A line of position obtained at one time may be used at a later time. All you need to do is move the line parallel to itself, a distance equal to the run of the ship in the interim, and in the same direction as the run. Such a line of position cannot be as accurate as a new line because the amount and direction of its movement can be determined only by the usual DR methods. If two new lines cannot be obtained, however, an old line, advanced and intersected with a new one, may be the only possible way of establishing a fix. Naturally, the distance an old line may be advanced without a substantial loss of accuracy depends on how closely the run can be reckoned.

In celestial navigation, as in piloting, you essentially are trying to establish the intersection of two or more lines of position. A single observation and the resulting LOP is insufficient to obtain a fix.

The most accurate method of obtaining a celestial fix is to take sights on many bodies in a short time. For example, it is quite common to take sights on six or more stars in a period of 15 minutes or less. Taking sights on many bodies allows the observer to identify and throw out LOPS with obvious errors.

Determining the LOP

You might be entitled to complain that much has been said concerning what an LOP tells you, but very little has been told about how you determine it in the first place. We are coming to that part now.

The first item is to take on a heavenly body or bodies and then reduce the sights. Reducing the sights taken gives you the information you need to plot the LOP. The LOPS then gives you the resulting fix.

The Celestial LOP

Determining the LOP, continued

Figure 9-8 illustrates the method used in establishing a single LOP by observing a star. An assumed position (AP) is selected according to certain requirements of convenience in calculating (described later). Observation of a star provides sextant altitude (h_s). Sextant altitude is then corrected to obtain observed altitude (H_o). The star's altitude from the assumed position, called the computed altitude (H_c), and its azimuth angle are determined from tables by a procedure you will soon learn. The azimuth angle is then converted to azimuth. After selecting an AP, draw the azimuth through the AP. Along the azimuth, measure off the altitude intercept (difference between the observed altitude and the computed altitude). At the end of this measurement, draw a perpendicular line, which is the LOP. You must know whether altitude intercept (a) should be measured from AP toward the body or from AP measured away from the body. It is helpful to remember the initials H_o MO T_o , if H_o is more toward. This means that if H_o is greater than H_c measure altitude intercept (a) from AP toward the body. If H_c is greater than H_o measure altitude intercept (b) from AP away from the body.

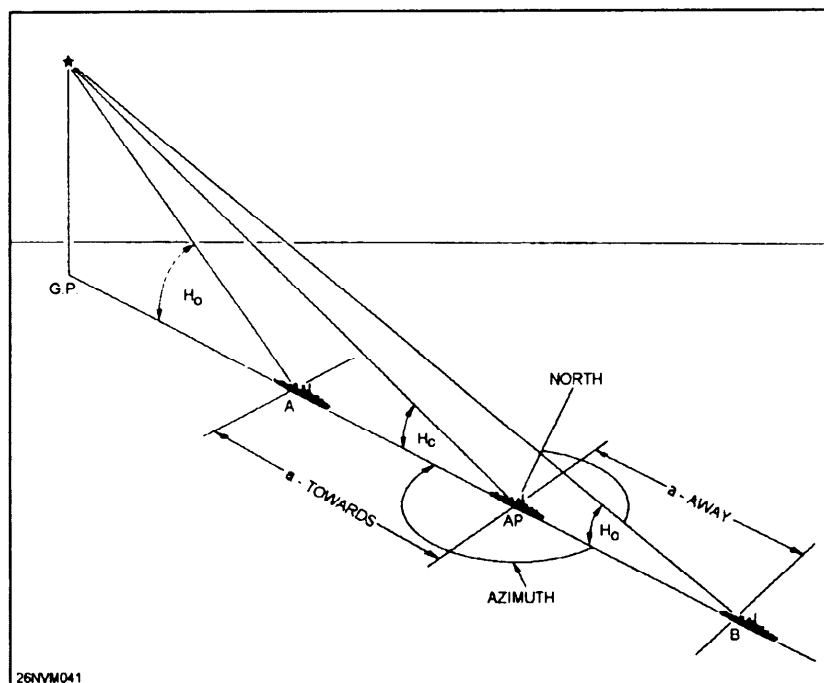


Figure 9-8. Determining a LOP.

Using the Sextant

General

The sextant is the instrument of chief importance in celestial navigation. It is used to measure the altitude of a heavenly body above the visible horizon. Sextant altitude is corrected for various factors to determine the body's true (or corrected) altitude above the celestial horizon.

Techniques

Here are some techniques commonly used to take sights with the marine sextant. It will always be necessary to find any index error prior to taking sights; refer to chapter 8 to find index error.

Use the following step action table for the general steps to take sights on the Sun. The steps for stars and planets are basically the same, except you would omit steps 2 and 4.

Step	Action						
1.	Hold the sextant level with the horizon and determine index error.						
2.	CAUTION: Set shade filters in place now, or eye burns may result.						
3.	Aim the sextant to a point on the horizon directly below the Sun.						
4.	<table><tr><th>IF...</th><th>THEN...</th></tr><tr><td>the Sun is rising</td><td>Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>below</i> the horizon.</td></tr><tr><td>the Sun is setting</td><td>Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>above</i> the horizon.</td></tr></table>	IF...	THEN...	the Sun is rising	Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>below</i> the horizon.	the Sun is setting	Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>above</i> the horizon.
IF...	THEN...						
the Sun is rising	Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>below</i> the horizon.						
the Sun is setting	Move the index arm slowly outward from the 0° position until the Sun's lower limb is just <i>above</i> the horizon.						
5.	Swing the arc. This means to gently move your hand grasping the sextant handle in a small upward arcing motion. Up to the left, then back to the right. You will see the reflected image of the Sun arc back and forth.						
6.	Give the recorder a standby to mark (marking the exact time of the sight). Continue swinging the arc while turning the micrometer drum slightly until the lower limb of the Sun touches the horizon. At that exact moment, <i>mark</i> the time of the sight and record the sextant altitude.						

Altitude Corrections

Altitude Corrections

Of the following five altitude corrections, the first three apply to observations of all celestial bodies. The last two corrections are applicable only when the observed body belongs to the solar system. Figure 9-9 illustrates the correction problem. To obtain the true altitude, you must correct the sextant altitude of any celestial body for:

1. **Index error**, which is the constant instrument error caused by a lack of perfect parallelism between the index mirror and horizon glass when the sextant is set at 0° .
2. **Refraction**, which is the deviation of rays of light from a straight line caused by Earth's atmosphere.
3. **Dip of the horizon**, which is the difference in direction between the visible and celestial horizons caused by the observer's height above the surface.

If the observed body belongs to the solar system, corrections must also be made for:

4. **Parallax**, which is caused by the proximity of bodies of the solar system to Earth, resulting in a difference in altitudes measured from the surface of Earth and from the center of Earth. Such an occurrence is not true of other heavenly bodies whose distance from Earth is considered infinite.
5. **Semidiameter**, which results from the nearness of bodies of the solar system, which makes it necessary to consider the observed bodies as appreciable size instead of as mere points of light; for example, stars. The sextant altitude of such a body is obtained by bringing its disk tangent to the horizon. Semidiameter correction must be applied to find the altitude of the center.

Altitude Corrections, Continued

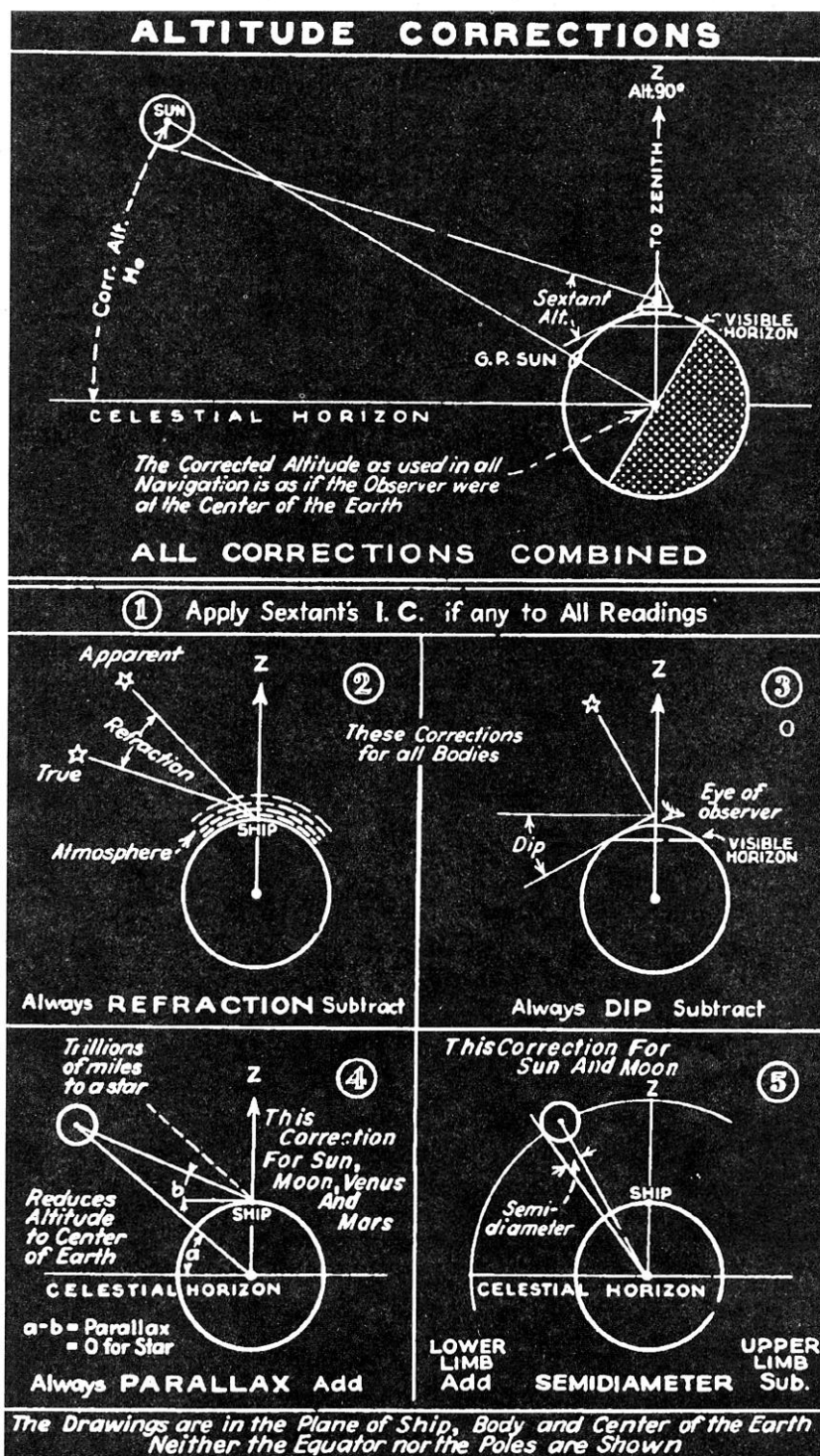


Figure 9-9. Sextant altitude corrections.

Altitude Corrections, Continued

Corrections Defined

We will explore each altitude correction in detail. Applying altitude corrections is the starting point for reducing sights for any observation.

Name	Description
Index Error	The amount of instrument error in the sextant (covered in chapter 8).
Refraction	<p>Earth is wrapped in a blanket of atmosphere more than 50 miles deep. Density of the atmosphere, like that of the ocean, increases with depth and is greatest at the bottom, next to Earth's surface. Light rays do not follow a straight line when passing through atmosphere of different densities, but are slightly bent into a gentle arc. This phenomenon is called refraction. Refraction is defined as the deviation of light rays from a straight line caused by their passage obliquely through mediums of different density. The measure of refraction is the angular difference between the apparent rays of light from an observed celestial body and its true direction.</p> <p>The effect of refraction is always to make the observed altitude greater than the true altitude. Consequently, refraction correction is always subtracted from the sextant altitude. Since refraction is caused by the oblique passage of rays through the atmosphere, rays from a body in the observer's zenith, intersecting the atmosphere at right angles, are not refracted. Maximum refraction occurs when a body is on the horizon, amounting then to between 34 and 39 minutes of arc. The amount of refractions depends on atmospheric conditions. Density of the atmosphere varies with barometric pressure and temperature. Refraction varies with density and also with the body's altitude. Because refraction varies with atmospheric conditions, and the effect of atmospheric conditions at low altitudes cannot be estimated with complete accuracy, observations of bodies below 10° should be regarded with suspicion. Refraction has no effect on the azimuth of a celestial body because it takes place entirely in the vertical plane of passage of the light rays.</p>
Dip	<p>The higher an observer's position is above the surface of the Earth, the more he/she must lower (or dip) the line of vision to see the horizon. Logically, then, all altitude observations must be corrected for the height of eye. Refer again to figure 9-9, and you will see why a dip correction is always subtracted.</p> <p>Failure to correct for dip from a height of 10 feet will result in an error of 3 miles in a line of position. From the bridge of the average destroyer, the resulting error would be approximately 10 miles.</p>

Altitude Corrections, Continued

Name	Description
Parallax	<p>Parallax is the difference between the altitude of a body, as measured from Earth's center, and its altitude (corrected for refraction and dip) as measured from Earth's surface. Altitude from the center of Earth is bound to be greater than from the surface. Consequently parallax is always a plus correction.</p> <p>Parallax increases from 0° for a body directly overhead to a maximum for a body on the horizon. In the latter instance, it is called horizontal parallax (HP). Parallax of the Moon is both extreme and varied because of its changing distance from Earth in its passage through its orbit. Parallax of the Sun is small; parallax of the planets is even smaller. For the stars, parallax is so tiny it is negligible.</p>
Semidiameter	<p>The true altitude of a body is measured to the center of that body. Because the Sun and Moon are of appreciable size, the usual practice is to observe the lower limb. Therefore, semidiameter correction must be added. It follows, then, that if the upper limb of either body is observed, the semidiameter correction is subtractive. Semidiameter correction amounts to about 16 minutes of arc for either the Sun or Moon. Stars are considered as points, and they require no semidiameter correction. When observing a planet, the center of the planet is visually estimated by the observer, so there is never a semidiameter correction.</p>

Remarks

In concluding the subject of altitude corrections, remember that **some tables** for altitude corrections (the *Nautical Almanac*, for example) combine two or more of the corrections for refraction, parallax, and semidiameter.

The correction for height of eye (dip) appears in a separate table for use with all bodies. Index error, which is impossible to include in such tables, should always be determined, recorded, marked plus or minus, and applied before any of the tabulated corrections.

Altitude Corrections, Continued

Strip Forms

The OPNAV Strip Form 3530/30; *H.O. 229*; *Nautical Almanac* are used to reduce sights for stars, planets, the Sun, and the Moon. The altitude corrections for each are the same except an additional correction is required for the Moon and planets. Reducing sights using this strip form is a process that can be broken down into the following stages:

Stage	Description
1.	Applying altitude corrections to find Ho (height observed).
2.	Using GMT to find LHA to enter Pub 229 with.
3.	Finding True Dec to enter Pub 229 with.
4.	Entering Pub 229 to find total corrections to apply to Ho to find Hc (height computed) Intercept, and Zn.

Steps to Follow to Find Ho

Use the following table to find Ho for any celestial body. Since we will be working several example problems, refer back to this table to find Ho.

Strip Form Pub 229 Naut Alm	Example problem to find Ho Action	Complete Strip Form Pub 229 Naut Alm
Body	Enter the symbol of the body.	SUN
GMT	Enter the GMT of the actual sight.	09 15 38
IC	Enter the value of the index correction.	- 1.0
D	Enter the dip correction (height of eye) from the inside cover of the <i>Nautical Almanac</i> .	- 6.9
Sum	Total the IC and D correction.	- 7.9
hs	Enter the uncorrected sextant altitude from the sight.	25° 46.9'
ha	Apply the sum to hs.	25° 39.0'
Alt Corr	Use ha to enter the altitude correction tables of the <i>Nautical Almanac</i> .	+ 14.3'
Add'l Corr Moon Hp/corr	Add any additional corrections for the Moon or planets.	N/A
Ho	Apply altitude and add'l corr to ha.	25° 53.3'

How to Reduce a Sunline Using Pub 229

Gather Information

As with any celestial observation, you must gather data to reduce to an LOP. With a sextant and recorder you will need the following: date/GMT of sight, DR position, sextant altitude (hs), height of eye of the observer, and IC correction.

Procedure

For our example we will use the following:

Date:	31 March 1984	GMT:	09 15 38
Lat:	36° 32.8'N	Long:	018° 10.0' W
hs:	25° 46.9'	IC:	- 1.0
Hgt of Eye:	50 ft		

After applying altitude corrections we have determined $H_o = 25^\circ 53.3'$.

We can now use the Pub 229 strip form to complete the process of reducing; at this point we have completed stage 1. We can move on to the next stage of finding LHA.

Notice that to find LHA, we follow the same steps as we did for our azimuth of the Sun problem only slightly different. Here's the key difference. We want to arrive at an even number LHA. To do this, we will use an assumed longitude. This step will help us in interpolation later in this problem. There is a catch though; the following rule must be adhered to when finding an assumed longitude.

Rule: The assumed longitude used as an assumed position must be within 30' of the original DR longitude.

Trick of the trade: When finding your assumed longitude, simply drop the minutes of total GHA down, then add the whole degree of longitude that is within 30' of the DR longitude. Look at our example problem where we dropped the 52.2' down from the total GHA. If we were to use the 18° from the original DR long. of 18° 10.0, which would be 18° 52.2' it would be more than 30', so we changed the 18° to 17° and all's well.

Let's begin working our problem on the next page.

How to Reduce a Sunline Using Pub 229, Continued

OPNAV 3130/30 Pub 229 Naut Alm	ACTION	Completed Strip Form
Ho	Apply altitude corrections to find.	25° 53.3'
GHA(h)	Enter the GHA hour value from the <i>Nautical Almanac</i> .	313° 57.7'
Increment (m/s)	Enter the minutes and seconds value from the <i>Nautical Almanac</i> .	3° 54.5'
Total GHA	Add GHA(h) and Increments (m/s).	317° 52.2'
v/v corr SHA	ENTER SHA for stars or planets only.	STARS and PLANETS ONLY
<i>a</i> Long (+E, -W) (+ - 360° if needed)	Enter the assumed DR longitude to arrive at an even degree of LHA, add east and subtract west.	17° 52.2'W
LHA	LHA= Total GHA + (v/v or SHA for star and planets) +E or -W DR Long.	300° 00.0'
Tab Dec	Enter the tabulated declination for 07 hours on the Sun column from <i>Nautical Almanac</i> .	N 4° 17.3'
d# / D Corr ⁿ	<p>The d# is found at the bottom of the Sun Dee column; in this case it is +1.0. It is assigned a + because dec is increasing.</p> <p>The D corr is found on the Increments and Corrections page for 15m38s. Look under the <i>v</i> or <i>d</i> column for the d# (1.0) and record the Corrⁿ value (0.4). The D Corrⁿ assumes the same sign as the d#.</p>	+0.6 / +0.3'
True Dec	Apply the D Corr ⁿ to Tab Dec	N 4° 17.6'
DR Lat same or contrary	Enter the whole degree of latitude and determine if it is named (N or S) as True Dec. In this case lat. is N and Dee is N, so it is same.	N 37° same

We have finished stages 2 and 3 and can move on to our final stage.

How to Reduce a Sunline Using Pub 229, Continued

OPNAV 3130/30 Pub 229 Naut Alm	ACTION	Completed Strip Form
Dec Inc /d	Dec Inc = True Dec min. only / d = d from Pub 229 entered with whole degrees of LHA, Dec, and Lat. (See fig. 9-10.)	17.6 / +38.1
Tens / DSD	Enter from the Pub 229 interpolation tables located on the inside of the front and back cover. (See fig. 9-11.)	+ 8.8
Units / DSD corr	Same as above.	+ 2.4
Total Corr	Total of tens and units.	+ 11.2
Hc (Tab)	Enter from Pub 229.	26° 07.5'
Hc (Comp)	Apply Total Corr to Hc (Tab).	26° 18.7'
Ho	Drop Ho down from the top of the form.	25° 53.3'
a	Subtract the higher value of either Hc(Comp) or Ho from the other. In this case, Ho is subtracted from Hc(Comp). The A means away. We will fully explain Towards and Away when we plot the LOP.	A 25.4
Z	Enter from Pub 229. Apply the rules for 2 just as with our azimuth problem.	105.8
Zn	LHA is greater than 180 so Zn = Z.	105.8

We have now completed the sight reduction solution for a sunline. The goal was to obtain an LOP. Where is the LOP you ask? Everything we need is right here. We will use the Zn (true bearing), a (intercept), and assumed position to plot our LOP. Let's move on to that task right now.

How to Reduce a Sunline Using Pub 229, Continued

60°, 300° L.H.A.

30°

32°

33°

34°

35°

36°

37°

Dec.

Hc

d

Z

Hc

d

Z

Hc

d

Z

Hc

d

Z

Hc

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Hc

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Z

Dec.

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25 39.5

26 12.6

26 45.4

27 17.8

27 49.7

28 21.3

28 52.5

29 23.3

29 53.6

30 23.5

30 52.9

31 21.9

-23.1

-22.8

-22.4

-21.9

-21.6

-21.3

-20.8

-20.3

-19.9

-19.4

-18.9

-18.4

106.1

105.2

104.2

103.3

102.3

101.4

100.4

99.4

98.4

97.4

96.4

95.4

25.3

25.5

26.0

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27.3

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28.4

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14.9

49.2

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-25.0

-34.6

-34.3

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-32.0

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107.0

106.1

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103.4

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101.5

100.5

99.6

98.6

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96.6

24 47.6

25 23.4

25 58.9

26 34.1

27 08.9

27 43.4

28 17.5

28 51.3

29 24.6

29 57.5

30 30.0

31 02.5

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60°, 300° L.H.A.

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60°, 300° L.H.A.

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Dec.

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-28.6

-29.2

-29.7

-30.3

-30.3

5.4

4.3

3.2

2.1

1.1

0.0

0.0

37 22.8

36 55.4

36 27.4

35 58.9

35 29.7

35 00.0

34 00.0

-27.4

-28.0

-28.5

-29.2

-29.7

-30.3

-30.3

5.5

4.3

3.2

2.2

1.1

0.0

0.0

38 22.5

37 55.2

37 27.4

36 58.8

36 29.7

36 00.0

35 00.0

-27.3

-27.8

-28.3

-29.1

-29.7

-30.3

-30.3

5.6

4.4

3.3

2.2

1.1

0.0

0.0

60°, 300° L.H.A.

30°

32°

33°

34°

35°

36°

37°

Dec.

Hc

d

Z

Hc

d

Z

Hc

d

Z

Hc

d

Z

Hc

d

Z

Hc

d

Z

Dec.

84

85

86

87

88

89

90

32 51.3

32 24.0

31 56.2

31 27.9

30 59.1

30 29.8

30 00.0

-27.8

-27.8

-28.2

-28.8

-29.3

-29.8

-30.2

5.1

5.1

4.1

3.0

2.0

1.0

0.0

33.2

33.2

32.7

31.9

31.1

30.3

30.0

23.5

55.9

27.7

59.0

29.8

00.0

00.0

-27.6

-28.2

-28.7

-29.2

-29.8

-30.2

-30.3

5.2

4.2

3.1

2.1

1.0

0.0

0.0

33 23.3

34 55.8

35 27.6

36 27.5

37 27.4

38 27.3

39 27.2

-27.5

-28.2

-28.6

-29.2

-29.7

-30.3

-30.3

5.3

4.2

3.2

2.1

1.0

0.0

0.0

36 23.0

35 55.6

35 27.5

34 58.9

34 29.7

34 00.0

33 00.0

-27.4

-28.1

-28.6

-29.2

-29.7

-30.3

-30.3

5.4

4.3

3.2

2.1

1.1

0.0

0.0

37 22.8

36 55.4

36 27.4

35 58.9

35 29.7

35 00.0

34 00.0

-27.4

-28.0

-28.5

-29.2

-29.7

-30.3

-30.3

5.5

4.3

3.2

2.2

1.1

0.0

0.0

38 22.5

37 55.2

37 27.4

36 58.8

36 29.7

36 00.0

35 00.0

-27.3

-27.8

-28.3

-29.1

-29.7

-30.3

-30.3

5.6

4.4

3.3

2.2

1.1

0.0

0.0

60°, 300° L.H.A.

30°

32°

33°

34°

35°

36°

37°

Dec.

Hc

d

Z

Hc

d

Z

H

Figure 9-10. Excerpt from Pub 229.

INTERPOLATION TABLE																	
Dec. Inc.	Altitude Difference (d)																Double Second Diff. and Corr.
	Tens					Decimals					Units						
	10'	20'	30'	40'	50'	0'	1'	2'	3'	4'	5'	6'	7'	8'	9'		
16.0	2.6	5.3	8.0	10.6	13.3	0	0.0	0.3	0.5	0.8	1.1	1.4	1.6	1.9	2.2	2.5	1.0 3.0 4.9 6.9 8.9 10.8 12.8 14.8 16.7 18.7 20.7 22.7 24.6 26.6 28.6 30.5 32.5 34.5
16.1	2.7	5.3	8.0	10.7	13.4	1	0.0	0.3	0.6	0.9	1.1	1.4	1.7	2.0	2.2	2.5	
16.2	2.7	5.4	8.1	10.8	13.5	2	0.1	0.3	0.6	0.9	1.2	1.4	1.7	2.0	2.3	2.5	
16.3	2.7	5.4	8.1	10.9	13.6	3	0.1	0.4	0.6	0.9	1.2	1.5	1.7	2.0	2.3	2.6	
16.4	2.7	5.5	8.2	10.9	13.7	4	0.1	0.4	0.7	0.9	1.2	1.5	1.8	2.0	2.3	2.6	6.9 8.9 10.8 12.8 14.8 16.7 18.7 20.7 22.7 24.6 26.6 28.6 30.5 32.5 34.5
16.5	2.8	5.5	8.3	11.0	13.8	5	0.1	0.4	0.7	1.0	1.2	1.5	1.8	2.1	2.3	2.6	
16.6	2.8	5.5	8.3	11.1	13.8	6	0.2	0.4	0.7	1.0	1.3	1.5	1.8	2.1	2.4	2.6	
16.7	2.8	5.6	8.4	11.2	13.9	7	0.2	0.5	0.7	1.0	1.3	1.6	1.8	2.1	2.4	2.7	
16.8	2.8	5.6	8.4	11.2	14.0	8	0.2	0.5	0.8	1.0	1.3	1.6	1.9	2.1	2.4	2.7	18.7 20.7 22.7 24.6 26.6 28.6 30.5 32.5 34.5
16.9	2.9	5.7	8.5	11.3	14.1	9	0.2	0.5	0.8	1.1	1.3	1.6	1.9	2.2	2.4	2.7	
17.0	2.8	5.6	8.5	11.3	14.1	0	0.0	0.3	0.6	0.9	1.2	1.5	1.7	2.0	2.3	2.6	
17.1	2.8	5.7	8.5	11.4	14.2	1	0.0	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	20.7 22.7 24.6 26.6 28.6 30.5 32.5 34.5
17.2	2.8	5.7	8.6	11.4	14.3	2	0.1	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	
17.3	2.9	5.8	8.6	11.5	14.4	3	0.1	0.4	0.7	1.0	1.3	1.5	1.8	2.1	2.4	2.7	
17.4	2.9	5.8	8.7	11.6	14.5	4	0.1	0.4	0.7	1.0	1.3	1.6	1.9	2.2	2.4	2.7	
17.5	2.9	5.8	8.8	11.7	14.6	5	0.1	0.4	0.7	1.0	1.3	1.6	1.9	2.2	2.5	2.8	22.7 24.6 26.6 28.6 30.5 32.5 34.5
17.6	2.9	5.9	8.8	11.7	14.7	6	0.2	0.5	0.8	1.0	1.3	1.6	1.9	2.2	2.5	2.8	
17.7	3.0	5.9	8.9	11.8	14.8	7	0.2	0.5	0.8	1.1	1.4	1.7	2.0	2.2	2.5	2.8	
17.8	3.0	6.0	8.9	11.9	14.9	8	0.2	0.5	0.8	1.1	1.4	1.7	2.0	2.3	2.6	2.9	
17.9	3.0	6.0	9.0	12.0	15.0	9	0.3	0.6	0.8	1.1	1.4	1.7	2.0	2.3	2.6	2.9	24.6 26.6 28.6 30.5 32.5 34.5
23.5	3.9	7.8	11.8	15.7	19.6	5	0.2	0.6	1.0	1.4	1.8	2.2	2.6	3.0	3.3	3.7	
23.6	3.9	7.9	11.8	15.7	19.7	6	0.2	0.6	1.0	1.4	1.8	2.2	2.6	3.0	3.4	3.8	
23.7	4.0	7.9	11.9	15.8	19.8	7	0.3	0.7	1.1	1.4	1.8	2.2	2.6	3.0	3.4	3.8	
23.8	4.0	8.0	11.9	15.9	19.9	8	0.3	0.7	1.1	1.5	1.9	2.3	2.7	3.1	3.4	3.8	
23.9	4.0	8.0	12.0	16.0	20.0	9	0.4	0.7	1.1	1.5	1.9	2.3	2.7	3.1	3.5	3.9	26.6 28.6 30.5 32.5 34.5
23.0	3.9	7.8	11.8	15.7	19.6	0	0.0	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	
23.1	3.9	7.9	11.8	15.7	19.7	1	0.0	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	
23.2	3.9	7.9	11.9	15.8	19.8	2	0.1	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	
23.3	3.9	8.0	11.9	15.9	19.9	3	0.1	0.4	0.7	1.0	1.3	1.5	1.8	2.1	2.4	2.7	28.6 30.5 32.5 34.5
23.4	3.9	8.0	12.0	16.0	20.0	4	0.1	0.4	0.7	1.0	1.3	1.6	1.9	2.2	2.4	2.7	
23.5	3.9	7.8	11.8	15.7	19.6	5	0.2	0.6	1.0	1.4	1.8	2.2	2.6	3.0	3.4	3.8	
23.6	3.9	7.9	11.8	15.7	19.7	6	0.2	0.6	1.0	1.4	1.8	2.2	2.6	3.0	3.4	3.8	
23.7	4.0	7.9	11.9	15.8	19.8	7	0.3	0.7	1.1	1.4	1.8	2.2	2.6	3.0	3.4	3.8	30.5 32.5 34.5
23.8	4.0	8.0	11.9	15.9	19.9	8	0.3	0.7	1.1	1.5	1.9	2.3	2.7	3.1	3.4	3.8	
23.9	4.0	8.0	12.0	16.0	20.0	9	0.4	0.7	1.1	1.5	1.9	2.3	2.7	3.1	3.5	3.9	
24.0	4.0	8.0	12.0	16.0	20.0	0	0.0	0.3	0.6	0.9	1.2	1.5	1.8	2.1	2.4	2.7	

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Figure 9-11. Interpolation table from the inside cover of Pub 229.

Plotting One or More LOPs

Procedure

Follow the steps in the table and refer to the accompanying figures to plot LOPs.

Step	Action
1.	Plot the AP (assumed position). This is the whole degree of latitude and the assumed longitude. In our example problem this would be Lat $37^{\circ} 00.0' N$ Long $017^{\circ} 52.2' W$.
2.	Lay off the azimuth line (Z_n) from the AP toward or away from the body, depending on whether the observed altitude (H_o) is greater or less than the computed altitude (H_c).
3.	Measure in the proper direction, along the azimuth line, the difference between the observed and the computed altitude in miles and tenths of miles. This is the value of a or intercept.
4.	Draw a line at the extremity of a , perpendicular (add 90° to Z_n) to the azimuth line. At the time of observation, this perpendicular line is the LOP.
5.	Label the LOP with the time of observation and the name of the observed body.

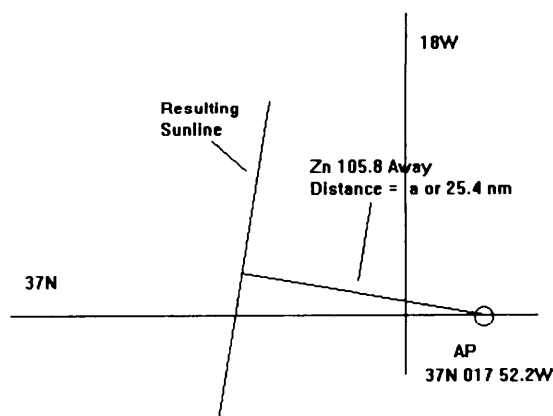


Figure 9-12. Plot the LOP.

Advancing LOPs

Advancing LOPs

Several methods may be used to advance a LOP. The most common method consists simply of advancing the AP in the direction of and for the distance of the run, as shown in figure 9-13, and drawing the new LOP.

Figure 9-13 illustrates a situation where the AP was advanced parallel to the course line for the distance run, and a new LOP was plotted from its new position. The new LOP was necessary because the same AP would have produced an LOP that would have intersected the course line beyond the limits of the chart. In this illustrative case, it is unnecessary to draw the first dashed construction on the chart.

The manner of advancing LOP from sights of the Moon, Venus, and Sirius (previously illustrated) to obtain an 1815 fix is seen in figure 9-14.

Advancing LOPs, Continued

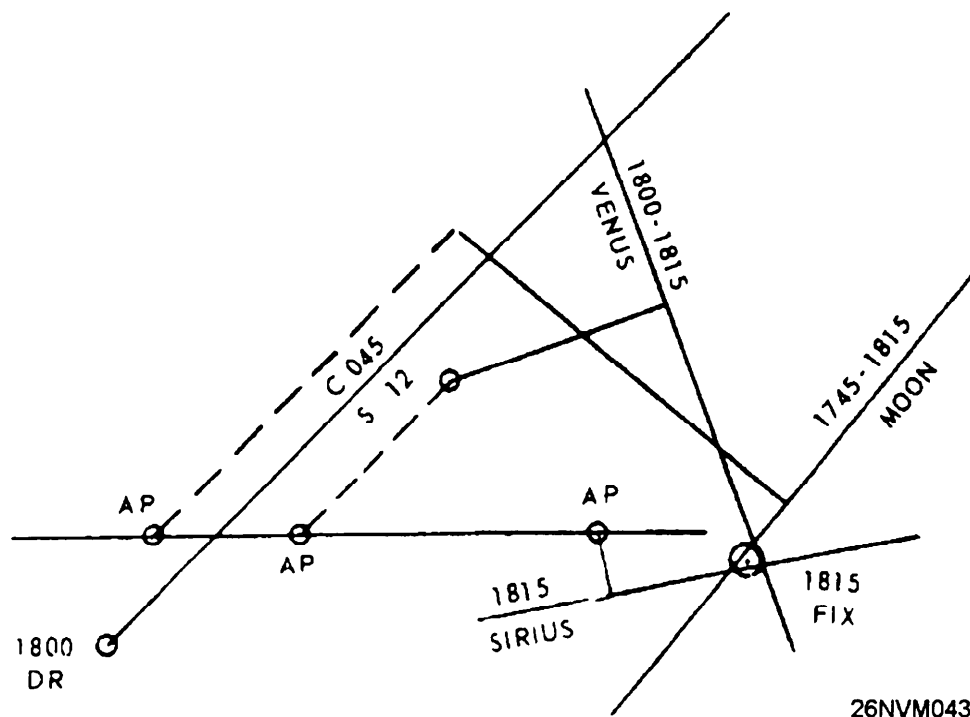


Figure 9-14. A fix from several LOPs.

Three lines of position by observation, like those obtained in piloting, do not always intersect exactly. Quite often a triangle is formed. If one or more of the LOPs must be advanced, the triangle is likely to be larger. Frequently, the center of the triangle is assumed to be the fix.

If, however, one or more lines have been advanced, more weight may be given to a line that has not been advanced, or to a line that the navigator has more confidence in; for example, favoring a first magnitude star over a third magnitude star. In figure 9-14, note that the plots are made from three separate APs, using the same assumed latitude but different assumed longitudes.

Reducing Stars, Planets, and the Moon Using Pub 229

Procedure for Stars and Planets

The steps to follow to reduce stars and planets are nearly the same steps that we used to reduce our sunline. The only differences are that in both cases when finding LHA, we must add the value of sidereal hour angle (SHA) to the total GHA to find LHA. Also, in the altitude correction tables for stars and planets there is an additional correction listed for some planets that must be added.

Procedure for the Moon

Once again the steps for the Moon are the same as our sunline except that the v and HP corrections must be added. These are additional altitude corrections. The v correction is always +. The HP correction for the nearest whole hour of GMT is selected. The v correction is found on the appropriate increments minutes and seconds page in the same manner as the d correction. If the upper limb of the Moon is observed, an additional correction (Add'l Corr) of -30' is made.

Reducing Sights Using Pub 249

General

HO Pub 249, *Sight Reduction Tables for Air Navigation* and the *Air Almanac*, can also be used to reduce sights. This method of sight reduction is used by some navigators; however, the degree of accuracy is slightly less than sight reduction by Pub 229.

Procedure

Once again, you would have to gather information to reduce; that is, GMT, sextant altitude, and so on.

For our example problem, we will use the following data:

Date: 30 March 1985 GMT: 06 26 21
 Lat: 36° 40.1'N Long: 017° 31.6' W
 hs: 40° 33.6' IC: + 0.8
 Hgt of Eye: 50 ft

OPNAV 3130/32 H.O.249 Air Alm	ACTION	Completed Strip Form
Body	Enter the name of the body.	REGULUS
GMT	Enter time of sight.	06h 26m 21s
IC	Enter the index correction.	+ 0.8
D	Enter the dip correction (hgt of eye 50ft) using the altitude correction table from the <i>Air Almanac</i> . (See fig. 9-16.)	- 7.0
R <i>o</i>	This is the refraction correction from the <i>Air Almanac</i> . (See fig. 9-17.)	- 1.0
SD		
ha		
Total Corr (sum)	Total of IC, D, and R <i>o</i> .	- 7.2
hs	Enter the sextant altitude.	40° 33.6'
Ho	Apply altitude corrections to find.	40° 26.4'

Reducing Sights Using Pub 249, Continued

CORRECTIONS TO BE APPLIED TO MARINE SEXTANT ALTITUDES									
CORRECTION FOR DIP OF THE HORIZON To be subtracted from sextant altitude									
Ht.	Dip	Ht.	Dip	Ht.	Dip	Ht.	Dip	Ht.	Dip
Ft.		Ft.		Ft.		Ft.		Ft.	
0		114	11	437	21	968	31	1 707	41
2	1	137	12	481	22	1 033	32	1 792	42
6	2	162	13	527	23	1 099	33	1 880	43
12	3	189	14	575	24	1 168	34	1 970	44
21	4	218	15	625	25	1 239	35	2 061	45
31	5	250	16	677	26	1 311	36	2 155	46
43	6	283	17	731	27	1 386	37	2 251	47
58	7	318	18	787	28	1 463	38	2 349	48
75	8	356	19	845	29	1 543	39	2 449	49
93	9	395	20	906	30	1 624	40	2 551	50
114	10	437		968		1 707		2 655	

26NVM048

Figure 9-16. Marine sextant altitude correction from the *Air Almanac*.

CORRECTIONS TO BE APPLIED TO SEXTANT ALTITUDE REFRACTION

To be subtracted from sextant altitude (referred to as observed altitude in A.P. 3270)

R _o	Height above sea level in units of 1000 ft.												R _o
	0	5	10	15	20	25	30	35	40	45	50	55	
Sextant Altitude													
	0	0	0	0	0	0	0	0	0	0	0	0	0
0	90	90	90	90	90	90	90	90	90	90	90	90	90
1	63	59	55	51	46	41	36	31	26	20	17	13	0
2	33	29	26	22	19	16	14	11	9	7	6	4	1
3	21	19	16	14	12	10	8	7	5	4	2	40	2
4	16	14	12	10	8	7	6	5	3 10	2 20	1 30	0 40	3
5	12	11	9	8	7	5	4 00	3 10	2 10	1 30	0 39	+0 05	4
6	10	9	7	5 50	4 50	3 50	3 10	2 20	1 30	0 49	+0 11	-0 19	5
7	8 10	6 50	5 50	4 50	4 00	3 00	2 20	1 50	1 10	0 24	-0 11	-0 38	6
8	6 50	5 50	5 00	4 00	3 10	2 30	1 50	1 20	0 38	+0 04	-0 28	-0 54	7
9	6 00	5 10	4 10	3 20	2 40	2 00	1 30	1 00	0 19	-0 13	-0 42	-1 08	8

26NVM049

Figure 9-17. Excerpt from refraction correction tables of the *Air Almanac*.

Reducing Sights Using Pub 249 , Continued

(DAY 089) GREENWICH A. M. 1985 MARCH 30 (SATURDAY)																			
GMT (UT)		☉ SUN		♈ ARIES		♀ VENUS-4.1		♃ JUPITER-2.1		♄ SATURN 0.3		☾ MOON		Lat.	Moon- rise	Diff.			
GHA Dec.		GHA °		GHA Dec.		GHA Dec.		GHA Dec.		GHA Dec.		GHA Dec.							
h m		° ' "		° ' "		° ' "		° ' "		° ' "		° ' "		N	o h m				
00	00	178 50.2	N 3 39.7	187 19.5		174 44 N14 10		234 09 S17 59		311 20 S17 24		82 38 N26 59							
10		181 20.3	39.9	189 49.9		177 14		236 39		313 51		85 03	59	72	☐	*			
20		183 50.3	40.1	192 20.3		179 45		239 10		316 21		87 27	59	70	☐	*			
30		186 20.3	40.2	194 50.7		182 16		241 40		318 52		89 52	59	68	☐	*			
40		188 50.4	40.4	197 21.1		184 46		244 10		321 22		92 16	58	66	☐	*			
50		191 20.4	40.5	199 51.5		187 17		246 41		323 52		94 40	58	64	☐	*			
05	00	253 51.2	N 3 44.6	262 31.8		250 03 N14 07		309 19 S17 59		26 33 S17 23		154 50 N26 50		56	16 15	19			
10		256 21.2	44.7	265 02.2		252 33		311 50		29 04		157 15	49	58	16 36	17			
20		258 51.2	44.9	267 32.6		255 04		314 20		31 34		159 39	49	60	17 03	14			
30		261 21.3	45.1	270 03.0		257 35		316 50		34 04		162 03	48						
40		263 51.3	45.2	272 33.5		260 05		319 21		36 35		164 28	48	S					
50		266 21.3	45.4	275 03.9		262 36		321 51		39 05		166 52	48						
06	00	268 51.4	N 3 45.6	277 34.3		265 06 N14 06		324 21 S17 59		41 36 S17 23		169 17 N26 47		Moon's P. in A.					
10		271 21.4	45.7	280 04.7		267 37		326 52		44 06		171 41	47	A	C	A			
20		273 51.4	45.9	282 35.1		270 08		329 22		46 36		174 05	46	I	T	I			
30		276 21.5	46.0	285 05.5		272 38		331 52		49 07		176 30	46						
40		278 51.5	46.2	287 35.9		275 09		334 23		51 37		178 54	46	+	+	+			
50		281 21.5	46.4	290 06.3		277 40		336 53		54 08		181 18	45	0	55	32			
07	00	283 51.6	N 3 46.5	292 36.7		280 10 N14 06		339 23 S17 59		56 38 S17 23		183 43 N26 45		5	56	56			
10		286 21.6	46.7	295 07.1		282 41		341 54		59 09		186 07	44	12	55	57			
20		288 51.6	46.9	297 37.6		285 11		344 24		61 39		188 32	44	16	58	30			
30		291 21.7	47.0	300 08.0		287 42		346 54		64 10		190 56	43	19	54	59			
40		293 51.7	47.2	302 38.4		290 13		349 25		66 40		193 20	43	22	53	28			
50		296 21.7	47.3	305 08.8		292 43		351 55		69 10		195 45	42	24	52	61			
11	00	343 52.3	N 3 50.4	352 46.6		340 25 N14 03		39 31 S17 58		116 48 S17 23		241 28 N26 33		53	33				
10		346 22.3	50.6	355 17.0		342 56		42 02		119 19		243 52	32	55	32				
20		348 52.4	50.7	357 47.4		345 27		44 32		121 49		246 16	32						
30		351 22.4	50.9	0 17.8		347 57		47 02		124 20		248 41	31						
40		353 52.4	51.1	2 48.2		350 28		49 33		126 50		251 05	31	Sun SD	16.0				
50		356 22.5	51.2	5 18.6		352 59		52 03		129 21		253 30	30	Moon SD	15'				
Rate 15 00.2 N0 01.0 15 03.8 S0 00.7 15 02.0 N0 00.1 15 02.6 0 00.0 14 26.3 S0 02.5														Age 9d					

Figure 9-18. Excerpt from the daily pages of the *Air Almanac*.

LAT 37°N

LHA	Hc	Zn	Hc	Zn	Hc	Zn	Hc	Zn	Hc	Zn	Hc	Zn	Hc	Zn	Hc	Zn
T	* VEGA	Alpheratz	ARCTURUS	* SPICA	REGULUS	* MIMULUS	CAPELLA									
180	16 15 053	44 14 087	55 52 112	38 11 154	54 27 233	36 10 280	18 48 314									
181	16 54 054	45 02 087	56 36 113	38 32 155	53 48 234	35 23 281	18 13 314									
182	17 33 054	45 49 088	57 20 114	38 52 156	53 09 236	34 36 281	17 39 315									
183	18 12 055	46 37 089	58 04 115	39 11 157	52 29 237	33 49 282	17 05 315									
184	18 51 055	47 25 089	58 47 116	39 29 158	51 49 238	33 02 282	16 32 316									
185	19 30 056	48 13 090	59 30 118	39 46 160	51 08 239	32 15 283	15 58 316									
186	20 10 056	49 01 090	60 12 119	40 02 161	50 27 240	31 28 283	15 25 317									
187	20 50 057	49 49 091	60 54 120	40 17 162	49 45 241	30 42 284	14 53 317									
188	21 30 057	50 37 092	61 35 121	40 32 163	49 03 242	29 55 284	14 20 318									
189	22 10 057	51 25 092	62 16 123	40 45 165	48 20 243	29 09 285	13 48 318									
190	22 51 058	52 13 093	62 56 124	40 57 166	47 37 244	28 23 285	13 16 318									
191	23 31 058	53 01 094	63 35 126	41 08 167	46 54 245	27 37 286	12 44 319									
192	24 12 059	53 48 094	64 13 127	41 18 169	46 10 246	26 51 286	12 13 319									
193	24 53 059	54 36 095	64 51 129	41 28 170	45 26 247	26 05 287	11 42 320									
194	25 34 059	55 24 096	65 28 131	41 36 171	44 42 248	25 19 287	11 11 320									
255	* DENEZ	ALTAIR	Nunki	* ANTARES	ARCTURUS	* Alkaid	Kocab									
256	48 32 061	42 31 116	21 18 153	26 11 188	49 49 256	53 17 306	49 39 347									
257	49 13 061	43 14 117	21 40 154	26 03 189	49 03 256	52 38 306	49 28 347									
258	49 55 061	43 57 118	22 00 155	25 55 190	48 16 257	51 59 306	49 17 347									
259	50 37 061	44 39 119	22 21 156	25 47 191	47 29 258	51 20 306	49 05 346									
260	51 19 061	45 21 120	22 40 157	25 37 192	46 42 259	50 41 306	48 54 346									
261	52 01 061	46 03 121	22 59 157	25 26 193	45 55 260	50 02 306	48 42 346									
262	52 44 062	46 44 122	23 17 158	25 15 194	45 08 260	49 23 306	48 30 345									
263	53 26 062	47 25 123	23 34 159	25 03 195	44 21 261	48 44 306	48 18 345									
264	54 08 062	48 05 124	23 51 160	24 50 196	43 33 262	48 05 306	48 05 345									
265	54 50 062	48 44 125	24 07 161	24 37 197	42 46 262	47 27 306	47 53 345									
266	55 32 062	49 24 126	24 22 162	24 22 198	41 58 263	46 48 306	47 40 344									
267	56 15 062	50 02 127	24 36 163	24 07 199	41 11 264	46 09 306	47 27 344									
268	56 57 062	50 40 128	24 50 164	23 52 200	40 23 264	45 30 306	47 14 344									
269	57 39 062	51 17 130	25 03 165	23 35 201	39 35 265	44 52 306	47 00 344									
270	58 22 062	51 54 131	25 15 166	23 18 202	38 47 266	44 13 307	46 46 343									

Figure 9-19. Excerpt from Pub 249, Volume I.

Reducing Sights Using Pub 249, Continued

OPNAV 3130/32 H.O.249 Air Alm	ACTION	Completed Strip Form
GHA(h)	Enter the GHA hour value of Aries from the <i>Air Almanac</i> to the nearest 10 minutes. (See fig. 9-18).	282° 35.1'
Increment (m/s)	Enter the minutes and seconds value from the interpolation tables of the <i>Air Almanac</i> .	1° 35.5'
Total GHA	Add GHA(h) and Increments (m/s).	284° 10.6'
+ - 360 (if needed)		
a Long (+E, -W)	Enter the assumed DR longitude to arrive at an even degree of LHA, add east and subtract west.	17° 10.6'W
LHA	LHA= Total GHA + (v/v or SHA for stars and planets) +E or -W DR Long.	267° 00.0'
a LAT	Enter the assumed latitude.	37°N
Hc	Enter Pub 249 volume I, with the whole degree of LHA for REGULUS, record Hc and Zn (last block). (See fig. 9-19.)	40° 23.0'
Ho	Drop down Ho from above.	40° 26.4'
a	Find the difference between Hc and Ho (remember to use Ho MO To). In this example, Ho is more than Hc, so it's named T for towards.	T 3.4'
Zn	Enter the Zn (true bearing).	264°

As you can see, using Pub 249 to determine a celestial LOP is a quick process compared to using Pub 229. Keep in mind that some amount of accuracy is lost.

Latitude by Local Apparent Noon (LAN)

Time of Meridian Passage

The purpose of knowing ahead of time the exact time of meridian passage (the Sun directly overhead) of the Sun is to allow the observer and recorder to arrive on the bridge a few minutes early. A latitude line from LAN is very useful. It is often used along with two morning sunlines to establish a noon celestial running fix. We will again be using a strip form to complete our sight reduction. First, we will find the time of meridian passage, then we will work the LAN solution. For our example problem, we will use the following data: Date: 30 March 84, DR Lat: 36°36.1'N, DR Long: 19° 22.3'W.

OPNAV 3130/35 LAN	ACTION		Completed Strip Form
DR Long	Enter the DR longitude.		19° 22.3'W
STD Meridian	Enter the standard meridian.		15
d long (arc)	Find the difference between STD Mer and DR Long.		4° 22.3'
d long (time)	Convert arc to time using the arc to time page in the <i>Nautical Almanac</i> .		+ 17 Min
LMT Mer Pass	From the daily pages in the <i>Nautical Almanac</i> for the given date, enter the time of meridian passage (bottom right of page).		1204
ZT LAN (1st est.)	IF...	THEN...	1221
	west of the standard meridian	Add d long (time) to LMT Mer Pass.	
	east of the standard meridian	Subtract d long (time) from LMT Mer Pass.	
Rev.DR Long	Enter revised DR.		19° 22.3'W
STD Meridian	Enter the standard meridian.		15
d long (arc)	Find the difference between STD Mer and DR Long.		4° 40.0'
d long (time)	Convert arc to time.		+19
LMT Mer Pass	Enter LMT for Mer Pass.		1204
ZT LAN (2nd est)	Add time to LMT Mer Pass.		1223

Latitude by Local Apparent Noon (LAN), Continued

Taking Sights to Observe LAN Up to this point we have learned how to find the time that the Sun should be directly overhead. Now we need to know how to observe LAN. We will discuss two methods. The first is called following to maximum altitude; the second is called *numerous sights*.

Following to Maximum Altitude The oldest method of determining meridian altitude of the Sun, and the one used most commonly, is known as following to maximum altitude. It is recommended because of its adaptability to various conditions, and because its use develops an insight into how the altitude varies near the time of apparent noon.

At approximately 10 minutes before watch time of LAN, the observer contacts the Sun's lower limb with the horizon in the sextant. He/she then swings the sextant from side to side, and adjusts it until the Sun, seen moving in an arc, just touches the horizon at the lowest part of the arc. This procedure is known as swinging the arc, which was described earlier in this chapter.

As the Sun continues rising, a widening space appears between its lower limb and the horizon. By turning the micrometer drum, the observer keeps this space closed and maintains the Sun in contact with the horizon. The change in altitude becomes slower and slower, until the Sun "hangs". While it is hanging, the observer swings the sextant to make certain of accurate contact with the horizon. He/she continues the observations until the Sun dips, which is a signal that the Sun is beginning to lose altitude. The sextant then shows the maximum altitude attained.

Latitude by Local Apparent Noon (LAN), Continued

Numerous Sights

The method of taking numerous sights is a modification of the maximum altitude method. It is useful under conditions where heavy seas, clouds, and the like may make steady observation impossible. Well before watch time of LAN, the observer begins taking a series of altitudes. Their number depends on the difficulties of the situation and the possible error in computed time of transit. He/she reads off the altitudes to a recording assistant, turning the tangent screw slightly after each observation to make sure that the next altitude is an independent sight. Observations are discontinued when the altitude definitely shows signs of decreasing.

Under favorable conditions, even a series of skillfully taken observations may show an occasional erratic deviation from the normal gradual rise and fall. After sights showing a radical difference from the preceding or succeeding series are discarded, however, the hang should become evident, and it should be possible to judge the maximum altitude. The figure selected will probably be less than the altitude shown in one observation and more than that below it. The result should give latitude with an error no more than 1'. This reading is considerably more accurate than could be obtained by a single sight under the conditions described.

Finding Latitude

As you now know, you must first obtain a sight of the Sun when it's at maximum altitude and the time of observation. With this and a DR position, we can reduce the sight to find latitude; now we can work the second part of our strip form.

Latitude by Local Apparent Noon (LAN), Continued

OPNAV 3130/35 LAN	ACTION	Completed Strip Form
LAT by LAN		
ZT LAN (obs)	Enter the ZT of the observation.	1221
ZD	Enter the zone description.	+ 1
GMT	Convert ZT to GMT.	1321
Tab Dec	Enter the tabulated declination for the Sun from the <i>Nautical Almanac</i> .	N 3° 57.9'
d# / d Corr	Enter the d# from the bottom of the Sun column, remember to find out if dec is + (increasing) or - (decreasing). Find the d Corr from the increments minutes and seconds pages for 21 minutes.	+ 1.0 / + 0.4
True Dec	Apply the d Corr to Tab Dec	N 3° 58.3'
IC	Enter the index correction.	+ 1.2
D	Enter the dip correction.	- 6.9
Sum	Enter the total of the IC and D.	- 5.7
Hs	Enter the uncorrected sextant altitude.	57 16.4
Ha	Apply the sum of the IC and dip corrections.	57 10.7
Alt Corr	Enter the altitude correction from the inside cover of the <i>Nautical Almanac</i> .	+ 15.6
Ho	Apply the Alt Corr to Ha.	57 26.3

Latitude by Local Apparent Noon (LAN), Continued

OPNAV 3130/35 LAN	ACTION	Completed Strip Form								
89° 60'	Enter 89° 60.0'.	89° 60.0								
HO (-)	Enter Ho.	57° 26.3								
Z Dist	Subtract Ho from 89 60.0.	32° 33.7'								
True Dec	Enter True Dec.	N 3° 58.3								
Lat	Use the following rules for making the declination correction:	36° 32.0'								
	<table><tr><th>IF...</th><th>THEN...</th></tr><tr><td>Lat and Dec are of different names</td><td>Lat = Z dist - Dec</td></tr><tr><td>Lat and Dec are of same names and Lat is less than Dec</td><td>Lat = Dec - Z dist</td></tr><tr><td>Lat and Dec are of same names and Lat is greater than Dec</td><td>Lat = Z dist + Dec</td></tr></table>		IF...	THEN...	Lat and Dec are of different names	Lat = Z dist - Dec	Lat and Dec are of same names and Lat is less than Dec	Lat = Dec - Z dist	Lat and Dec are of same names and Lat is greater than Dec	Lat = Z dist + Dec
	IF...		THEN...							
	Lat and Dec are of different names		Lat = Z dist - Dec							
	Lat and Dec are of same names and Lat is less than Dec		Lat = Dec - Z dist							
	Lat and Dec are of same names and Lat is greater than Dec		Lat = Z dist + Dec							
Time		12h 21m 00s								

Closing Remarks

Celestial navigation requires skill gained through experience. This chapter has given you the basic knowledge required to meet the minimum requirements of the Quartermaster occupational standards. This is just the tip of the iceberg; you should strive to perfect your celestial skills. In the event of a large scale war, you may find that all electronic means of obtaining a fix have been knocked out. It's important that electronic fixes are compared to celestial fixes whenever possible. Remember, the prudent navigator uses all available means to accurately fix the ship's position along the intended track.

Quartermasters should study sources other than this RTM to gain additional knowledge on celestial navigation. *Dutton's Navigation and Piloting* is a excellent reference on this material.